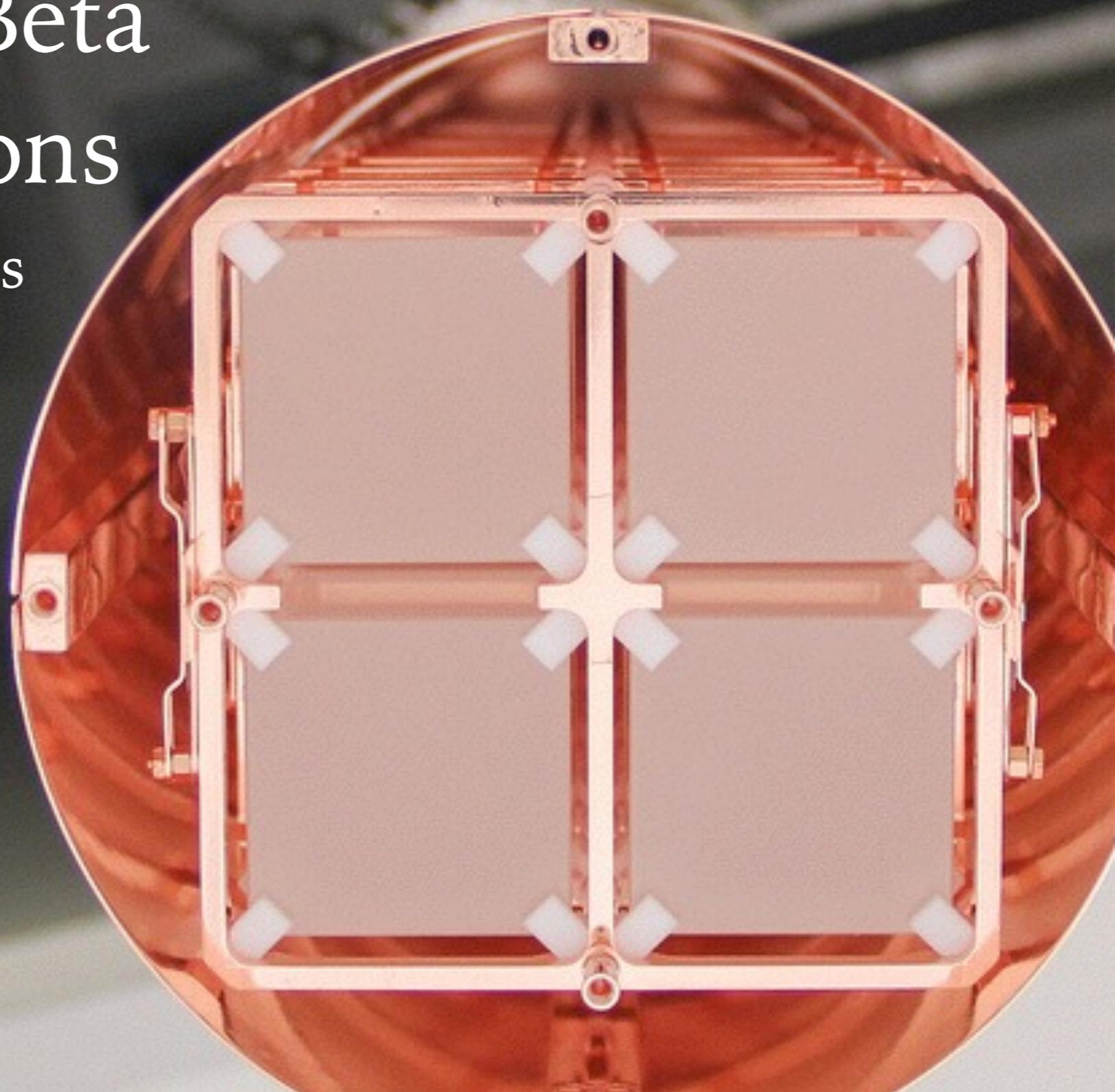


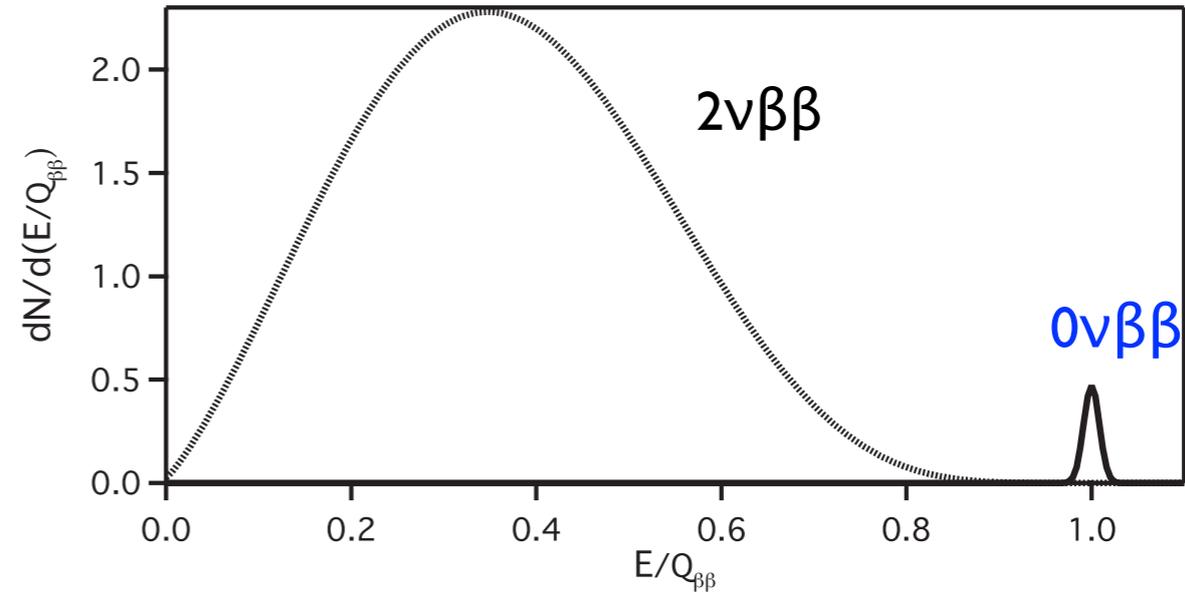
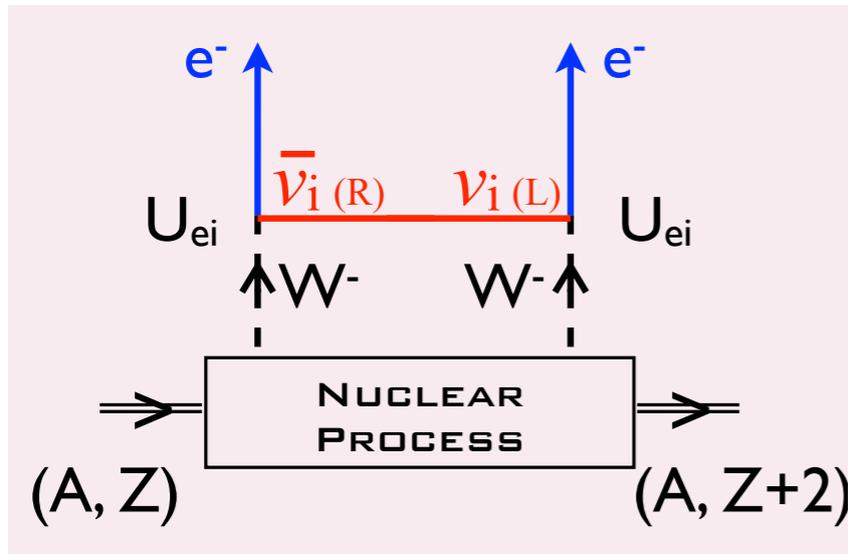
TES for Double Beta Decay Applications

Raul Hennings Yeomans
UC Berkeley



Neutrino-less Double Beta Decay

- ▶ Hypothetical $\beta\beta$ decay mode allowed if neutrinos are Majorana particles, i.e. $\bar{\nu}_i \equiv \nu_i$



Phase space factor Nuclear matrix element

$$\frac{1}{T_{1/2}^{0\nu}} = G^{0\nu} |M^{0\nu}|^2 |\langle m_{\beta\beta} \rangle|^2$$

Decay half-life Effective Majorana ν mass:

$$m_{\beta\beta} \equiv \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$$

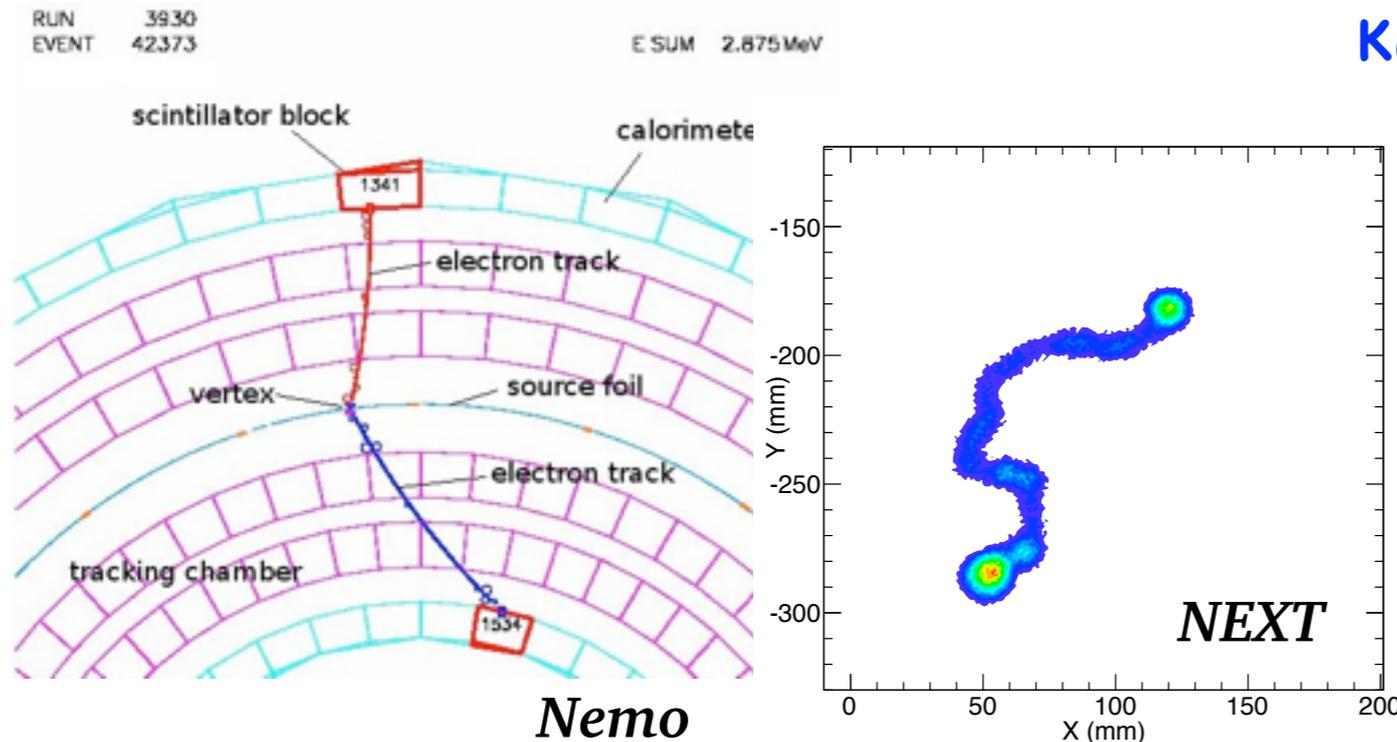
- ▶ For $m_{\beta\beta} = 15$ meV estimated half lives $10^{27} - 10^{28}$ years, depending on the nuclear system

- ▶ **Observation of $0\nu\beta\beta$ would mean**
 - Lepton number violation
 - Neutrinos are Majorana particles
 - Rate measures (effective) electron neutrino mass

Experimental approaches to $0\nu\beta\beta$

Source external to detector

Example: SuperNEMO



Plus: event topology, background rejection, multiple isotopes possible.

Cons: detector mass, resolution, acceptance.

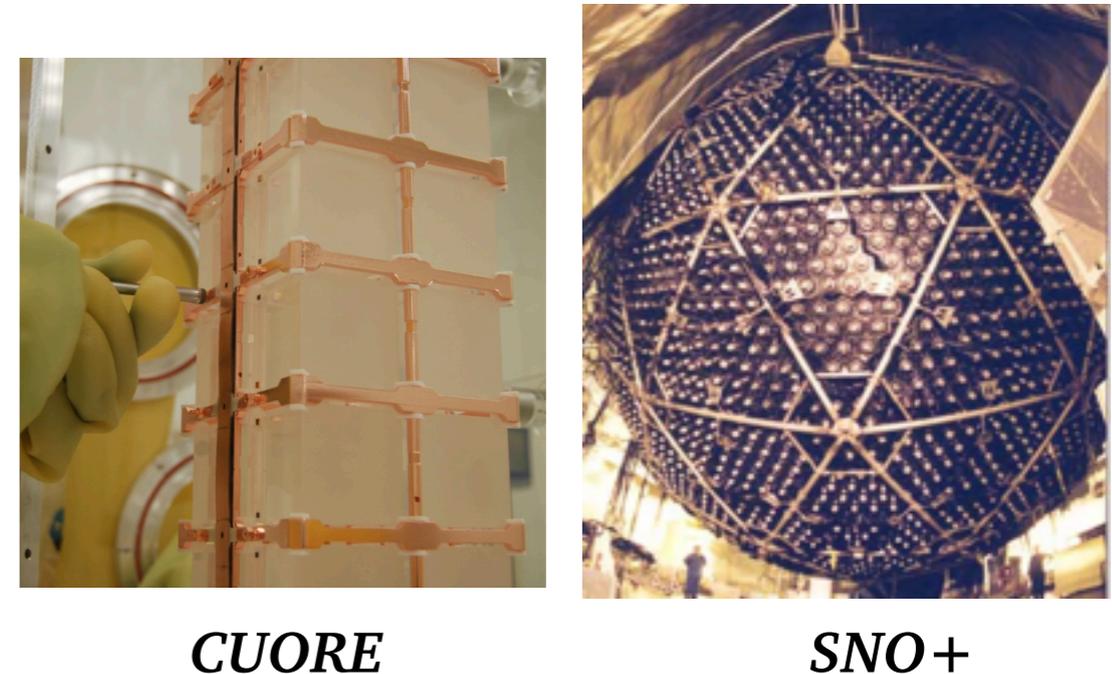
Technology: typically tracking detectors.

Technology:
Pressurized TPC
(10-16 atm)

May prove invaluable to test models
once $0\nu\beta\beta$ is discovered

Source internal to detector

Example: MAJORANA, EXO, CUORE, SNO+, Kamland-Zen, etc.



Plus: detector mass, energy resolution, acceptance

Cons: event topology, background rejection

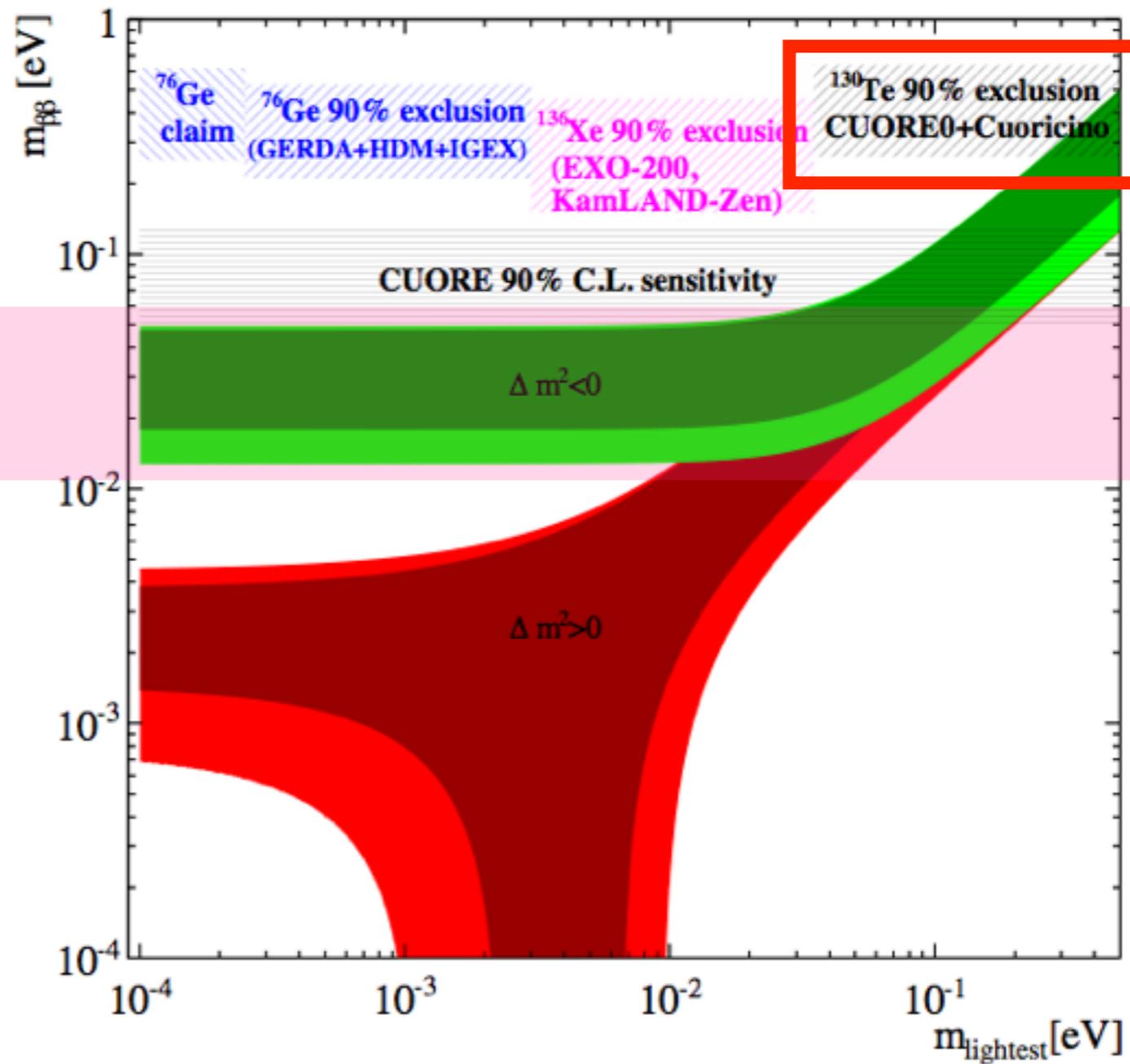
Technology: calorimeters (bolometers, ionization, scintillation), tracking

Typically aimed at $0\nu\beta\beta$ discovery

Beyond CUORE: towards covering the IHE

Goal of next generation experiments

$$m_{\beta\beta} \sim \frac{m_e}{\sqrt{F_N \cdot \epsilon \cdot \eta} \sqrt{\frac{M \cdot t}{b \cdot \delta E}}}$$



For a zero background experiment

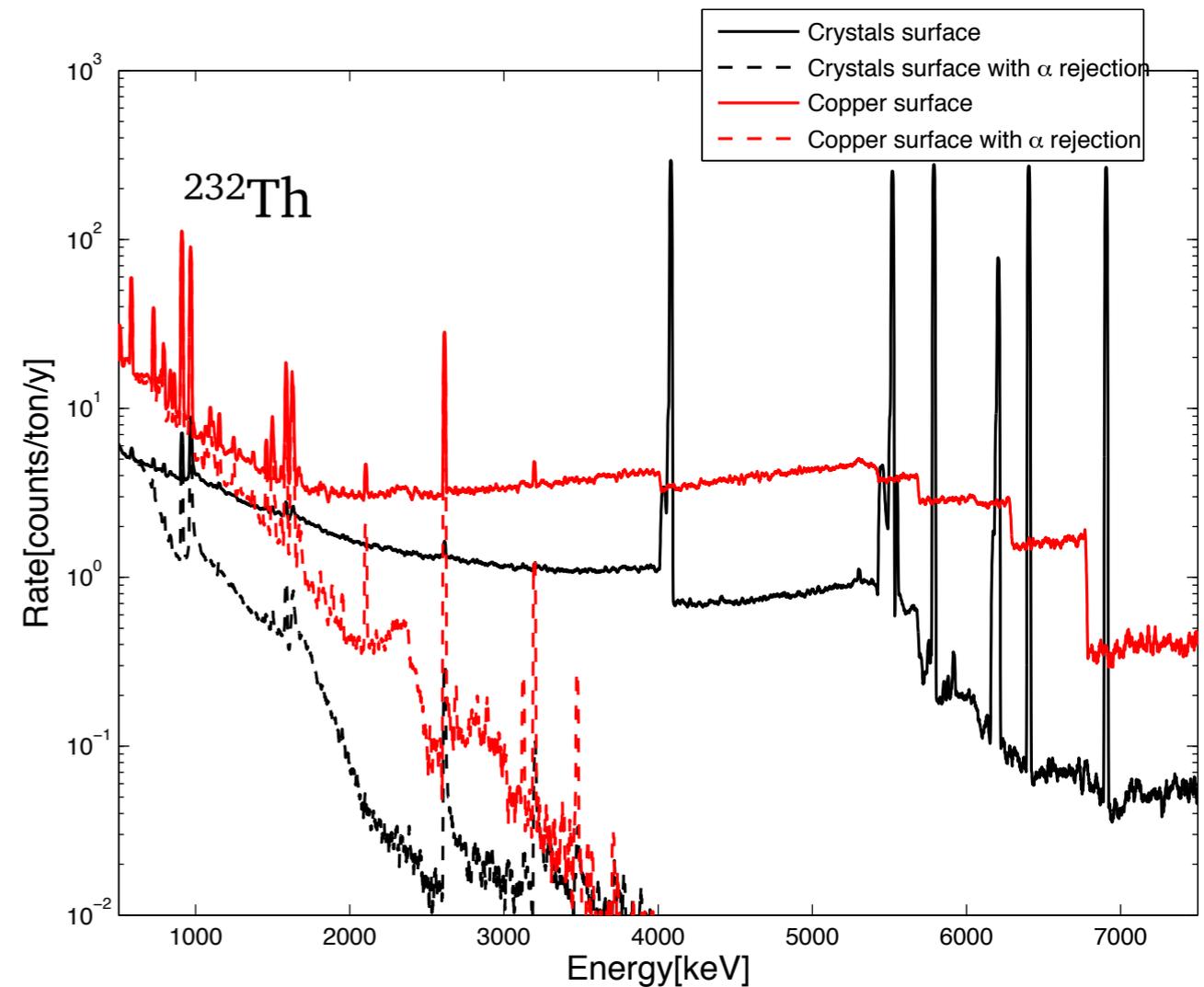
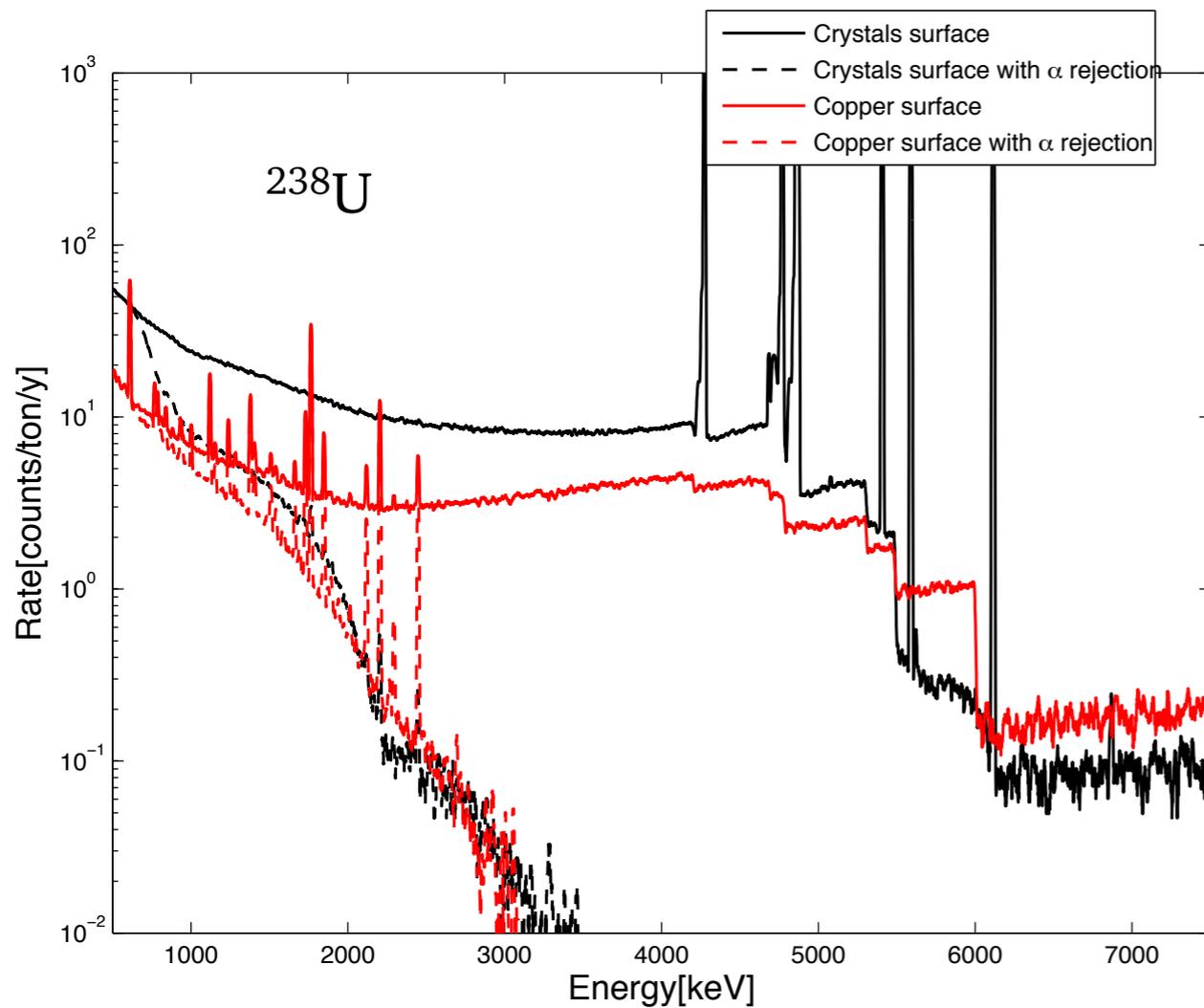
$$\langle m_\nu \rangle \propto 1 / \sqrt{T_{1/2}^{0\nu\beta\beta}} \propto 1 / \sqrt{Nt}$$

With background subtraction

$$\langle m_\nu \rangle \propto 1 / \sqrt{T_{1/2}^{0\nu\beta\beta}} \propto 1 / (Nt)^{1/4}$$

Beyond CUORE: effect of alpha background rejection

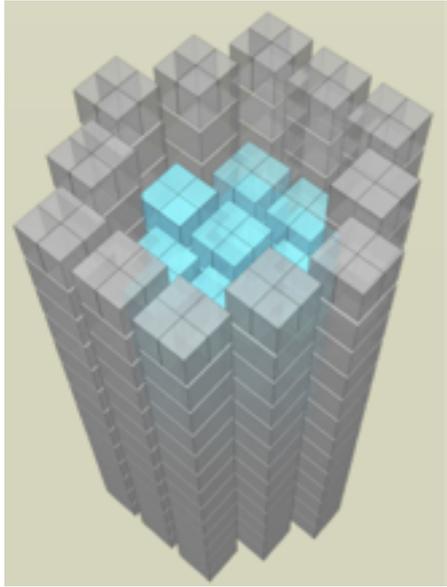
Simulation of surface contamination with an exponential depth profile and a mean depth of 5 μm . Dashed histograms are without α 's



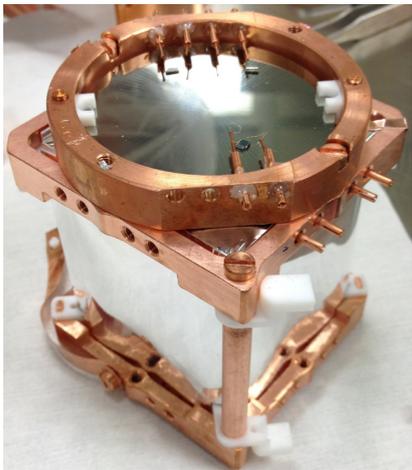
CUORE Collaboration Eur.Phys.J. C74 (2014) 10, 3096

CUPID: Cuore Upgrade with Particle ID

Te-130 enrichment

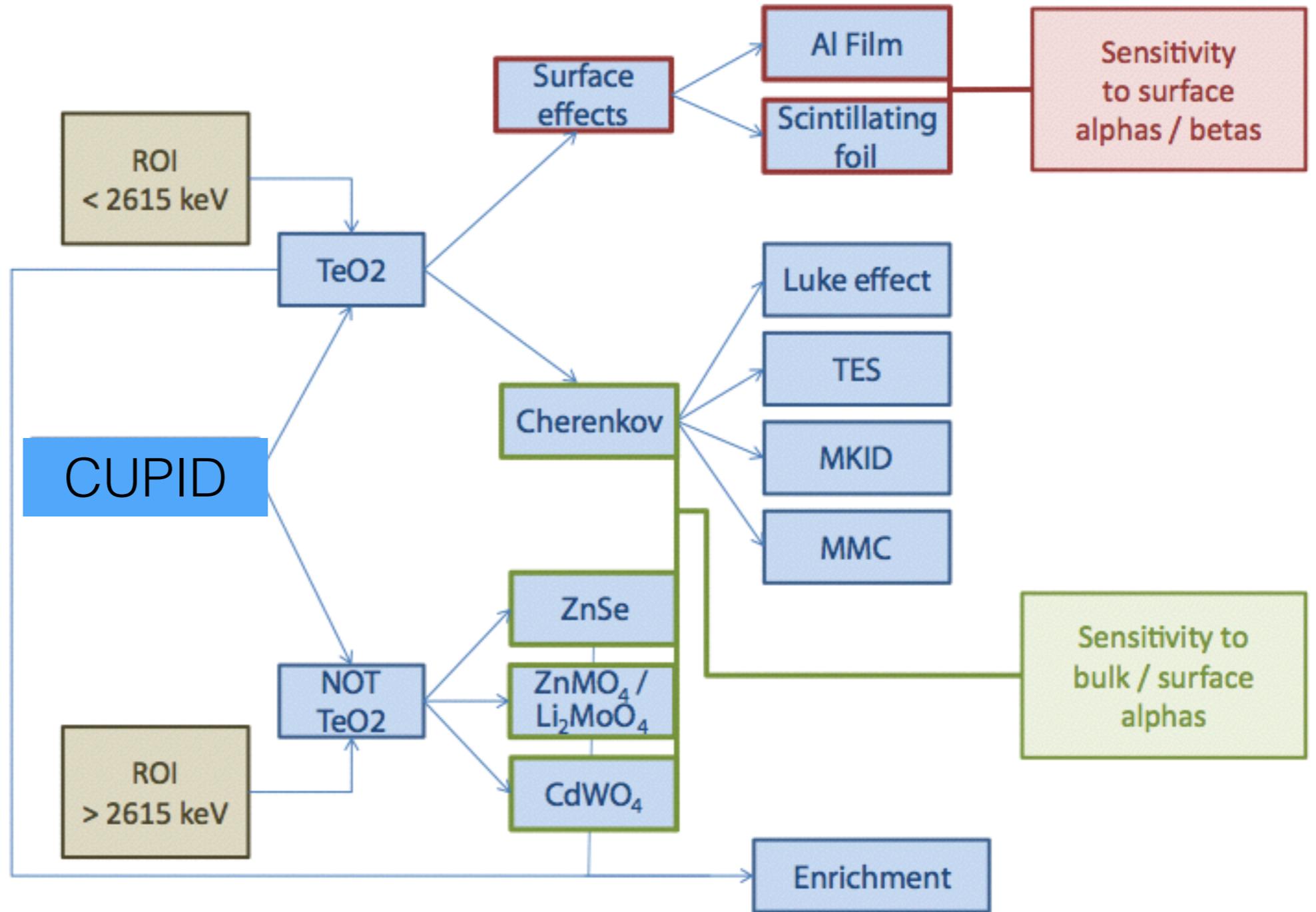


Particle ID



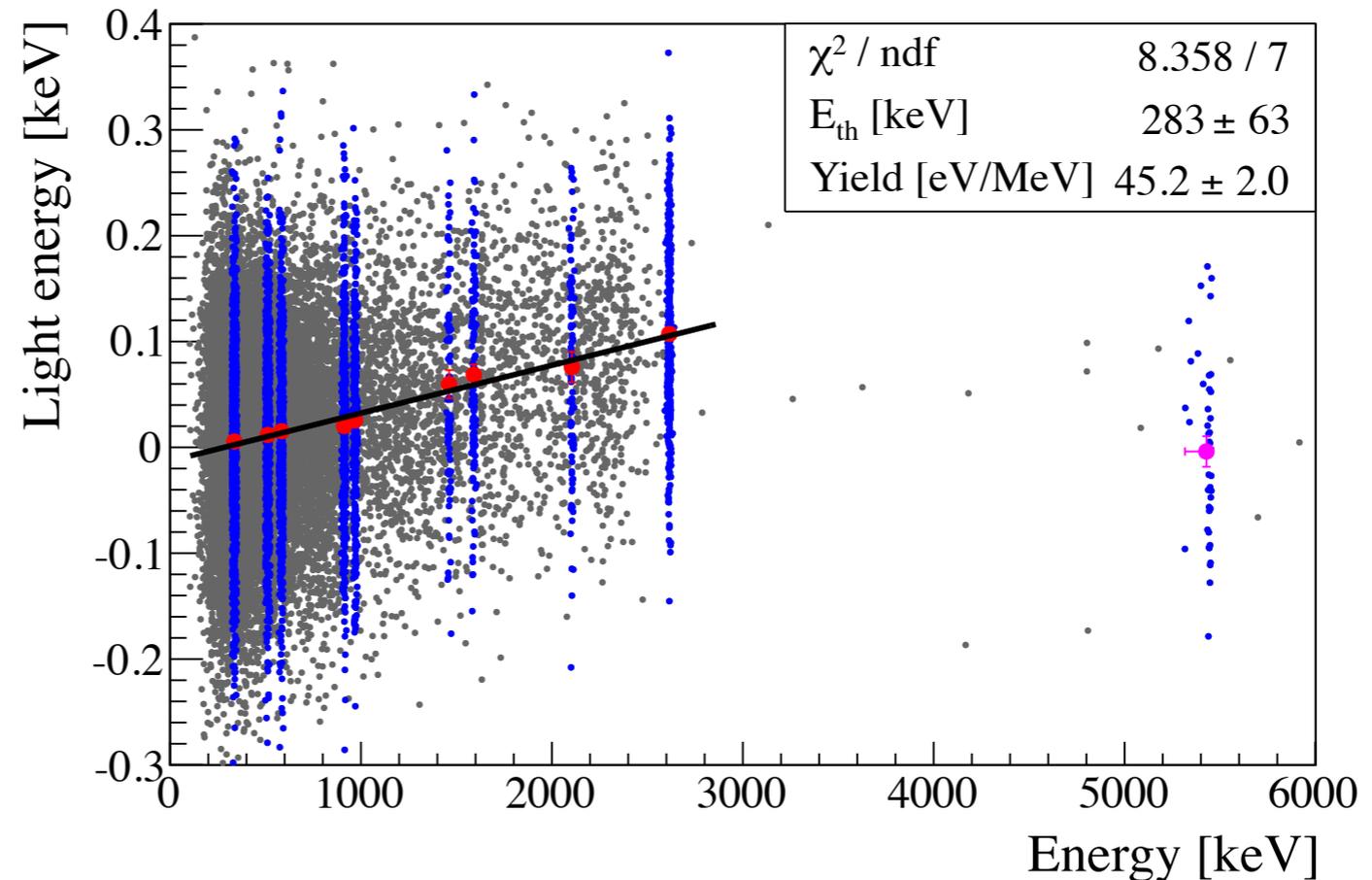
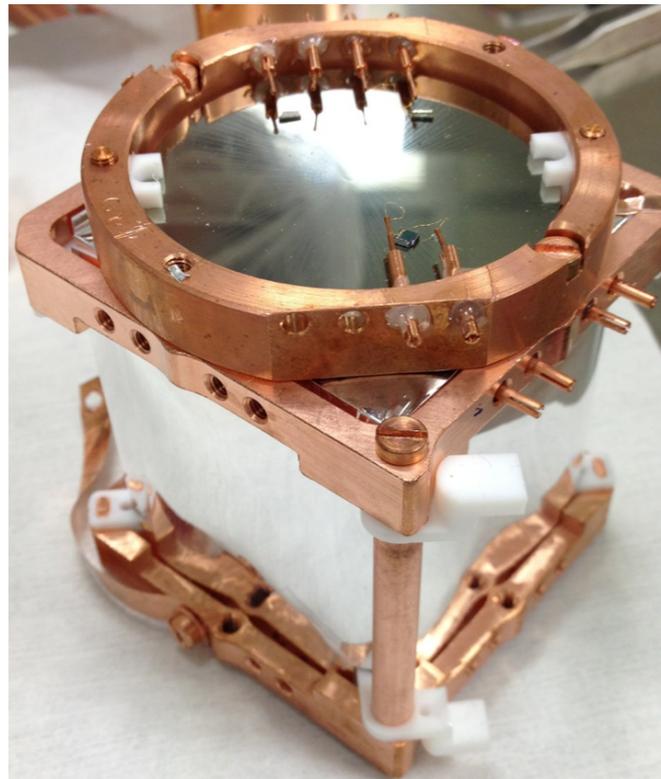
Bolometer R&D:

- CALDER
- Cherenkov/TeO₂
- LUCIFER
- LUMINEU



Particle ID via a the Light channel

Cherenkov light from a full 750g TeO₂ crystal



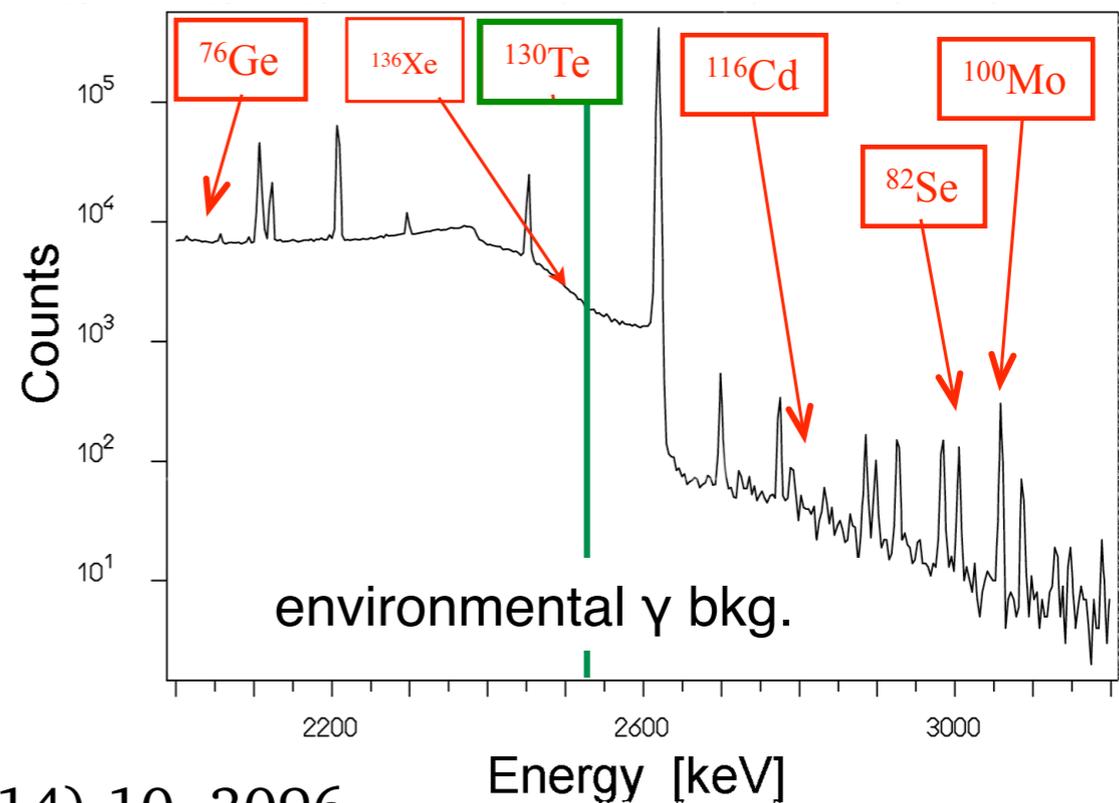
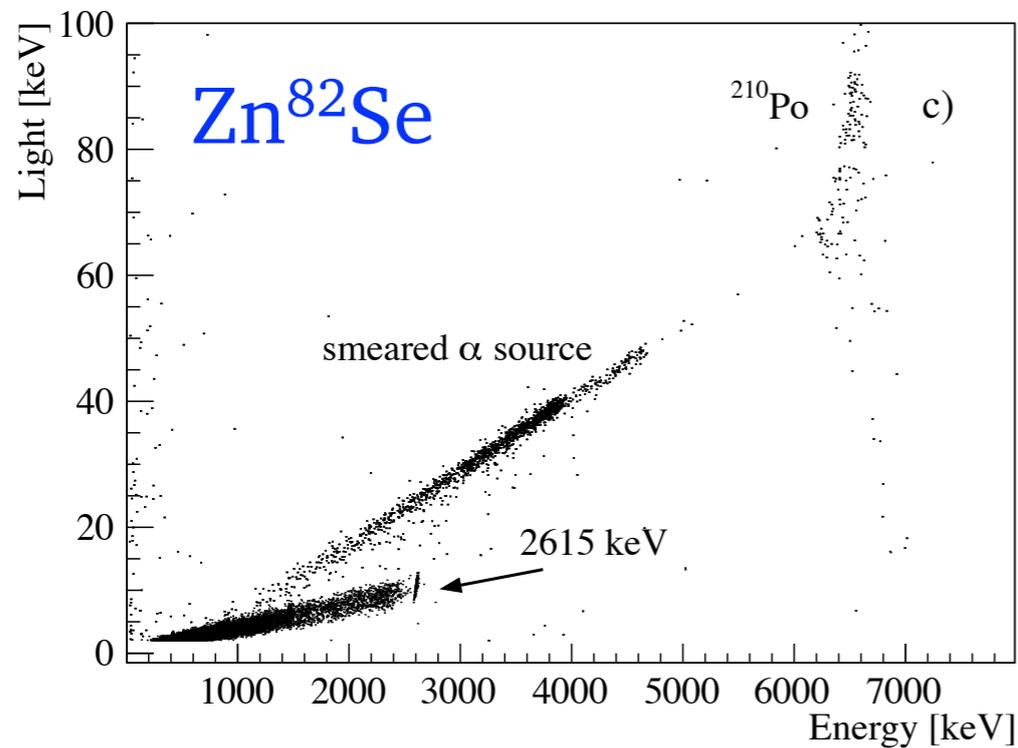
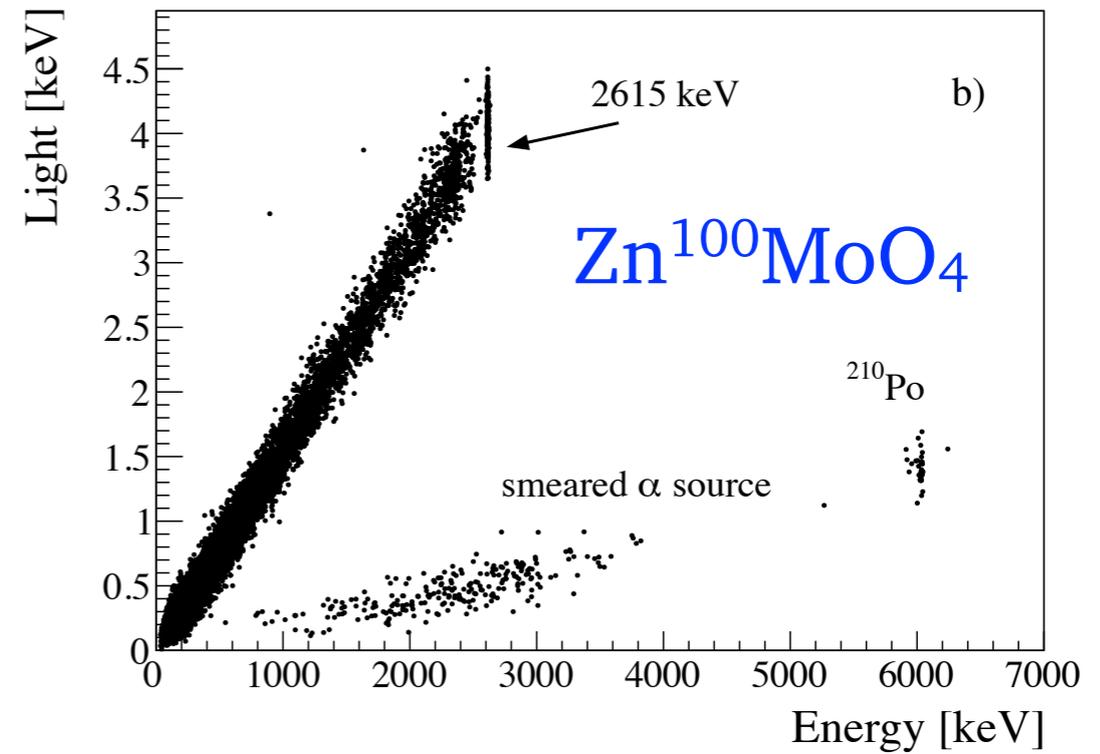
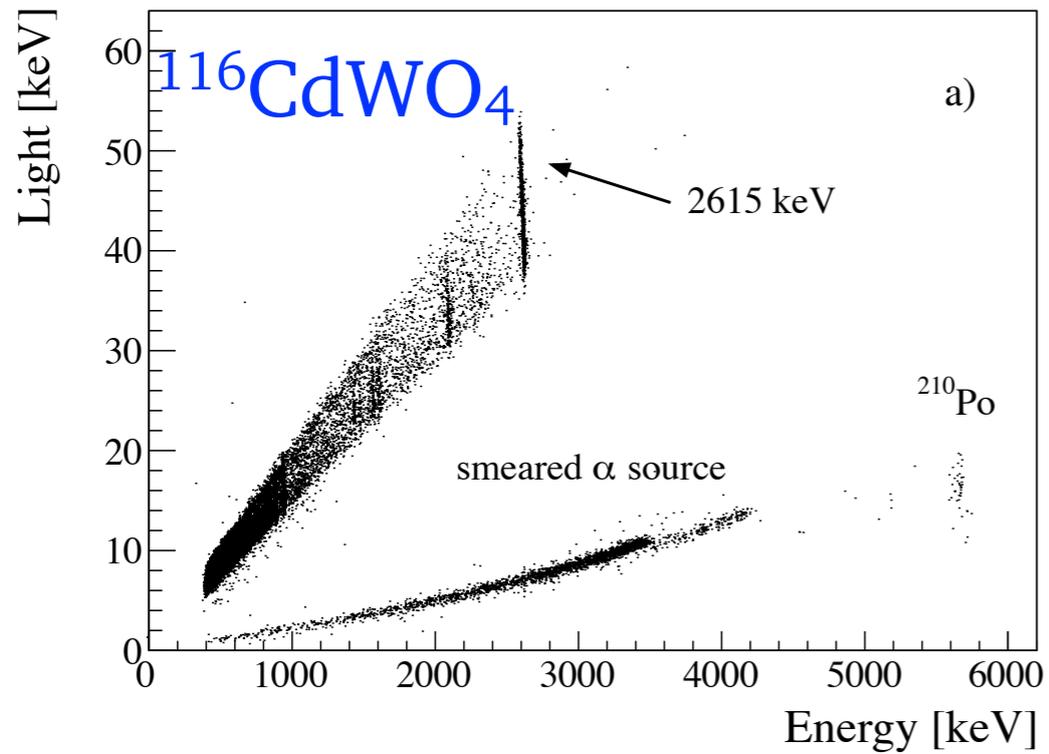
Casali et al. Eur. Phys. J. C75 (2015), 1, 12

101+/- 3.4 eV of light for a β/γ event with energy $0\nu\beta\beta$ value

R&D on improved light detectors currently ongoing:
MKIDs, TES, Neganov-Luke assisted, MMCs

We need to discriminate between α and β/γ at 5σ , ie 99.9% rejection α 's with 90% efficiency.
For a light yield of 100 eV implies resolution need of better than 20eV.

Scintillating crystals



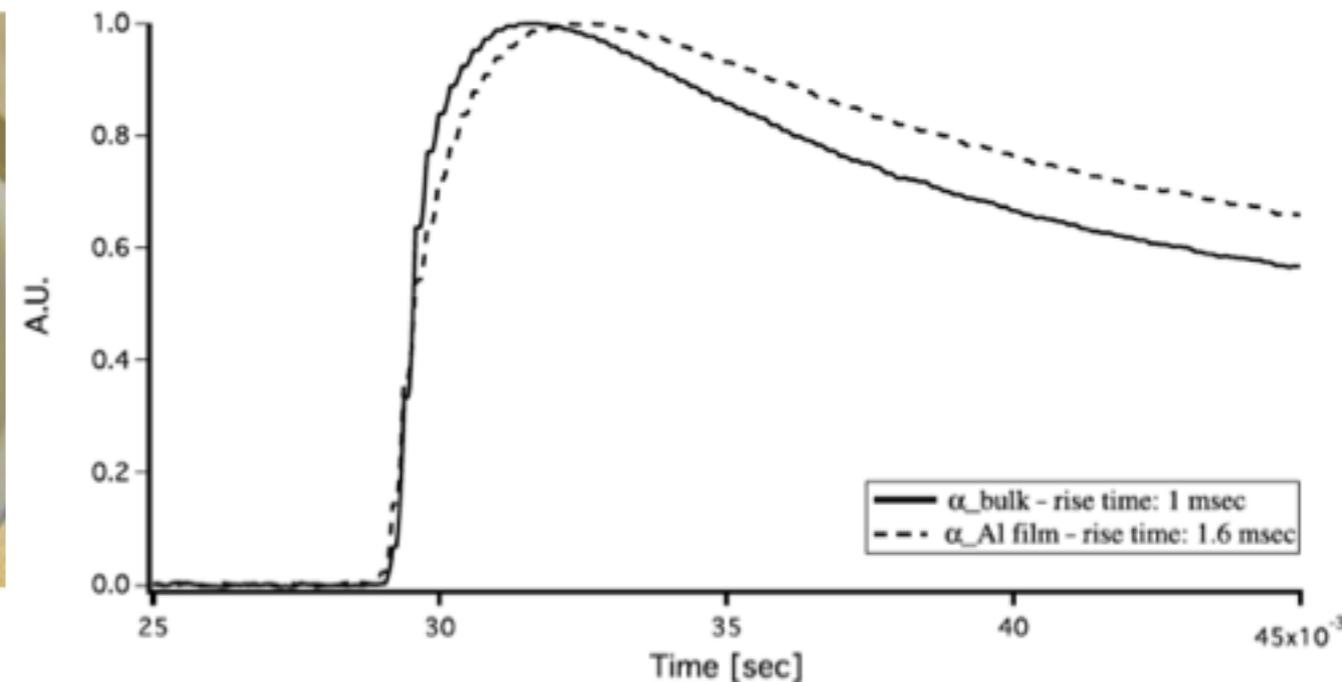
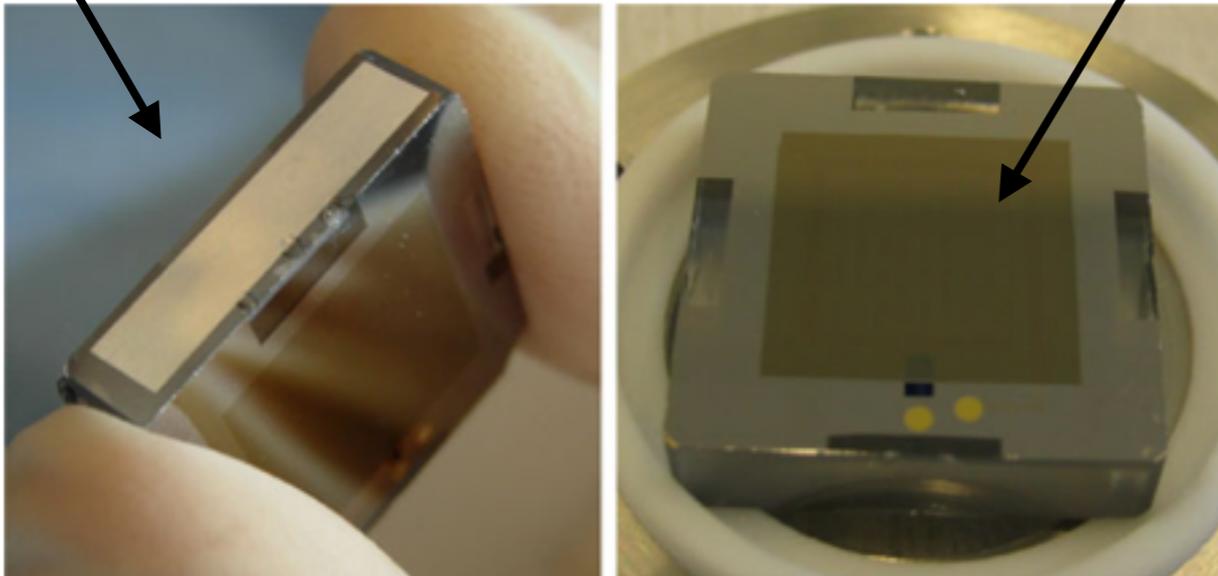
Cuore Collab. Eur. Phys. J. C74 (2014) 10, 3096

Advances in HEP, Vol 2013 (2013), Article ID 237973

Particle ID via Pulse Shape with Aluminum coating

Aluminum film

NbSi sensor



J. Low Temp. Phys (2012) 167:1029-1034

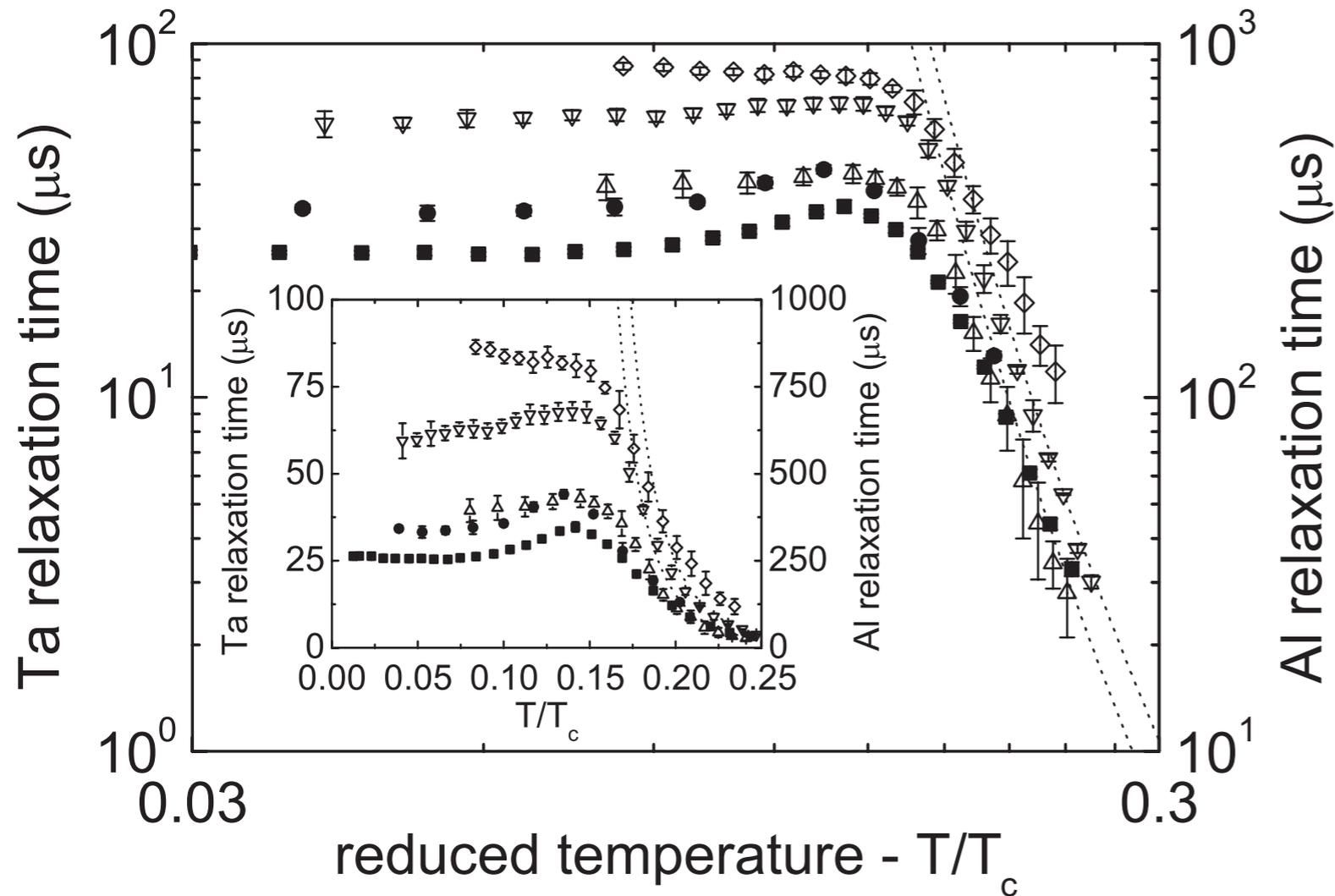
Quasiparticles “recombination” time:

$$\frac{1}{\tau_{\text{rec}}} = \frac{1}{\tau_0} \sqrt{\pi} \left(\frac{2\Delta}{kT_c} \right)^{5/2} \sqrt{\frac{T}{T_c}} e^{-\Delta/kT}$$

for example, for 1 μm of Al on sapphire $\tau_{\text{rec}} \sim 1.5 \mu\text{sec}$ from J. Shnagl NIMA 444 (2000)245-248

Quasiparticles recombination time τ_R

$$\frac{1}{\tau_{\text{rec}}} = \frac{1}{\tau_0} \sqrt{\pi} \left(\frac{2\Delta}{kT_c} \right)^{5/2} \sqrt{\frac{T}{T_c}} e^{-\Delta/kT}$$



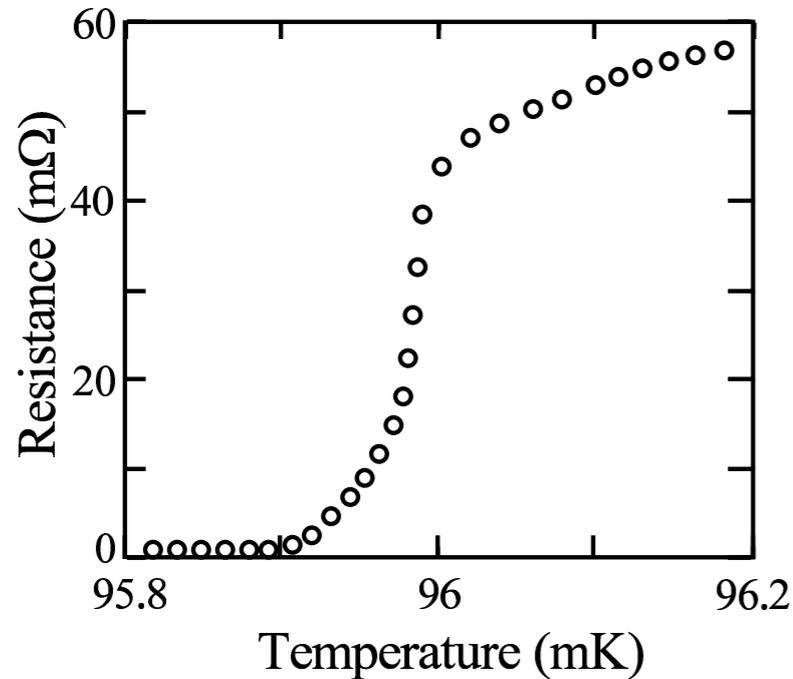
Film (substrate)	Thickness [nm]	τ_R [μsec]
Al(Si)*	100	390
Al(Si)*	250	600
Al(Al ₂ O ₃)*	250	860
Al(Al ₂ O ₃ **)	1000	1530 \pm 350

$\tau_R \sim \sqrt{\text{thickness}}$
 τ_R independent of Temp
 for $T/T_c < 0.1$
 τ_R depends on substrate

* R. Barends et al. PRL 100, 257002(2008)

** J. Schnagl NIMA 444 (2000)245-248

Transition Edge Sensors

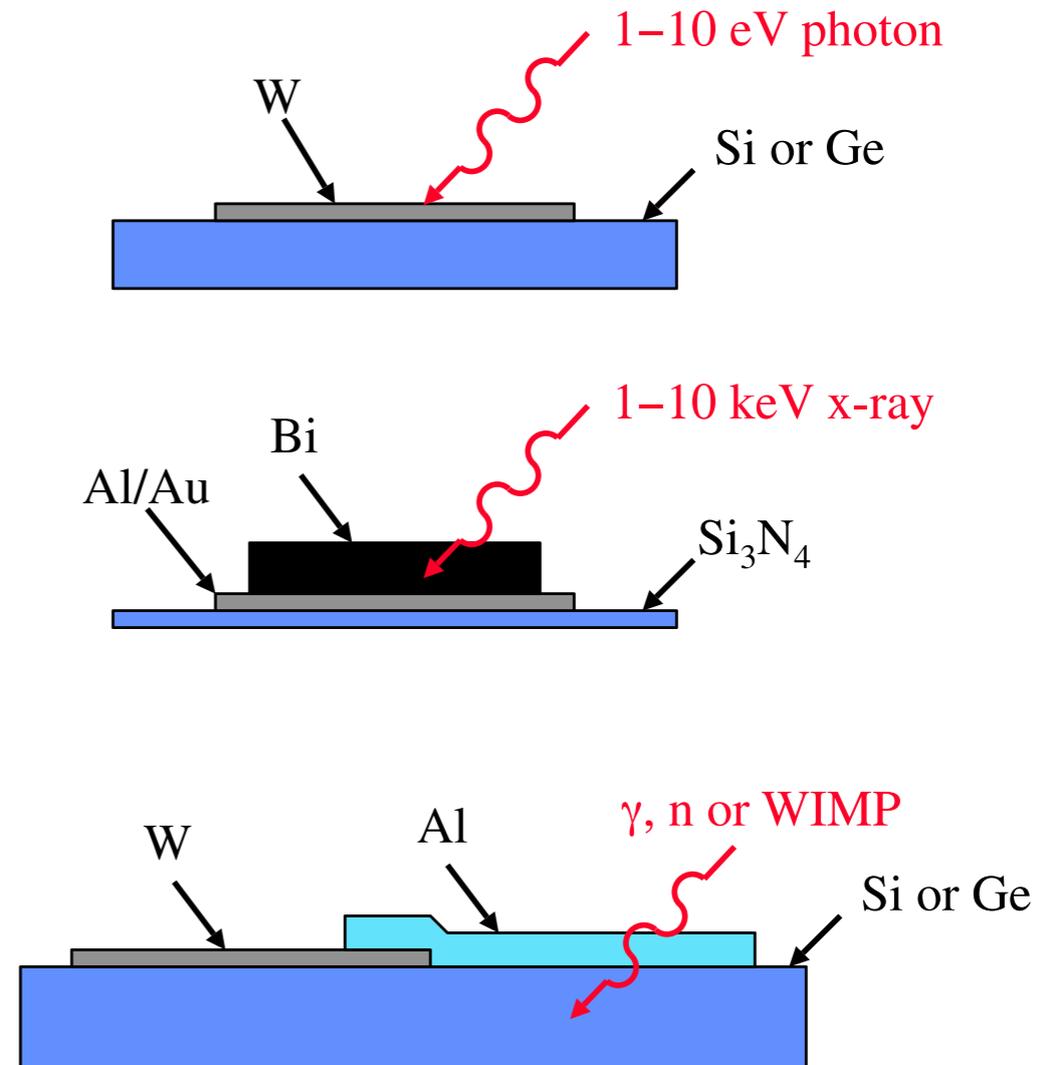


Topics Appl. Phys. 99, 63-149 (2005)

$$\Delta E_{\text{rms}} = \sigma_E \approx \sqrt{\frac{4k_B T^2 C_{\text{tot}}}{\alpha}} \sqrt{\frac{\beta + 1}{2}},$$

$$\alpha \equiv \frac{T}{R} \frac{dR}{dT}$$

β : determined by thermal conductivity between the TES and absorber/heat-bath

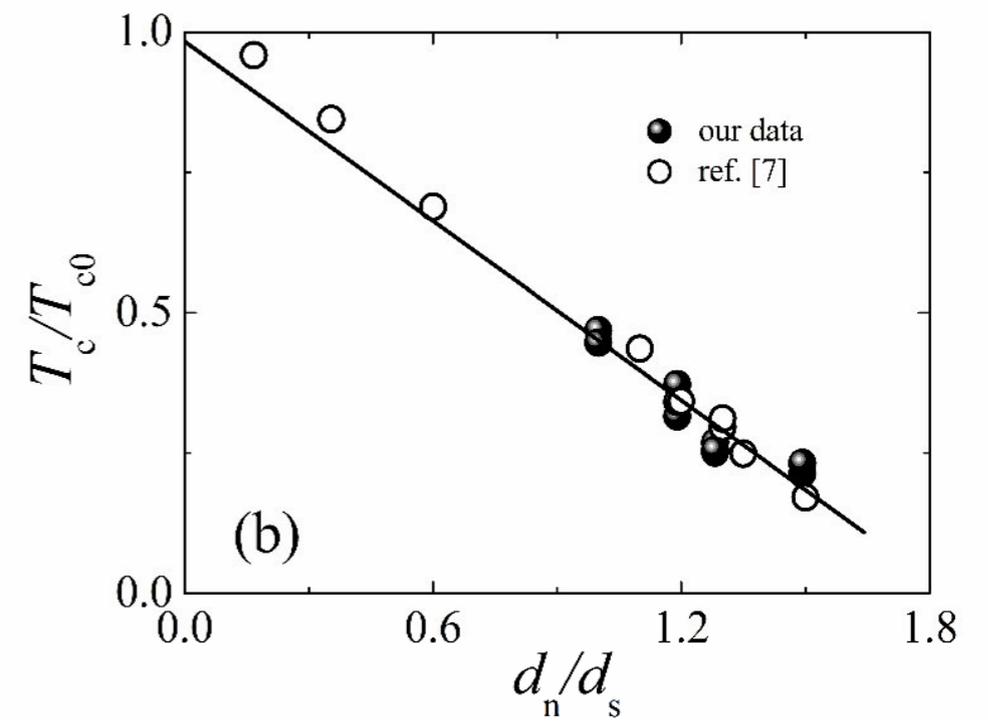
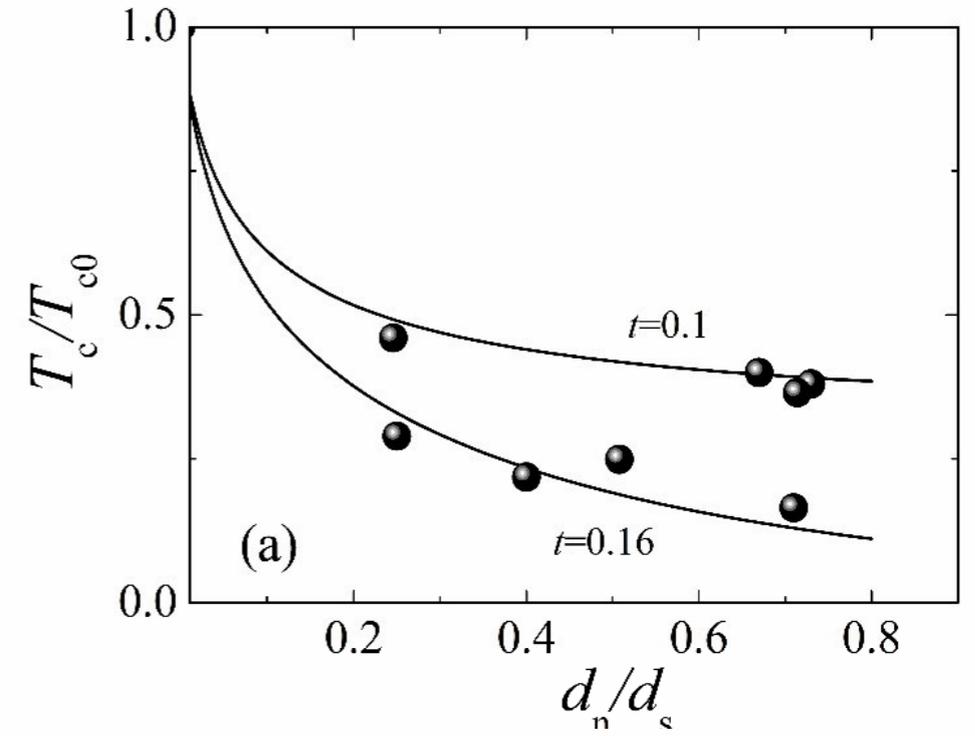
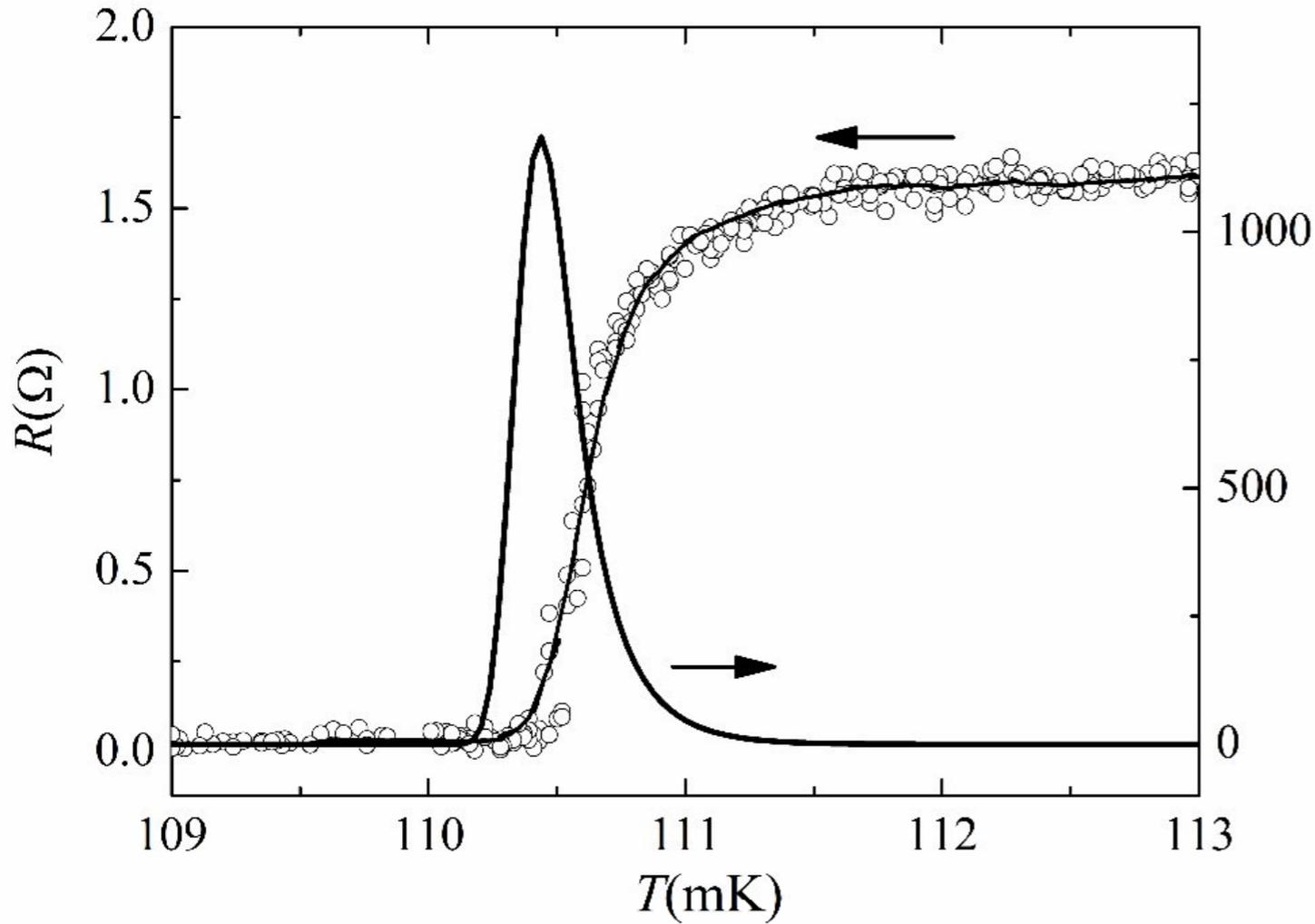


$$C_{\text{tot}} = C_{\text{bolo}} (\sim T^3) + C_{\text{TES}} (\sim T) + C_{\text{other}} \text{ (e.g. caused by impurities in the crystal)}$$

K.D. Irwin, Appl. Phys. Lett. Vol. 66, 1998 (1995)

Example of Ti/Pd bilayer TES

Journal of Physics: Conference Series **150** (2009) 052168



$$\Delta E_{\text{rms}} = \sigma_E \approx \sqrt{\frac{4k_B T^2 C_{\text{tot}}}{\alpha}} \sqrt{\frac{\beta + 1}{2}}$$

K.D. Irwin, Appl. Phys. Lett. Vol. 66, 1998 (1995)

TES scalability and low Tc



Tc tuning through $^{56}\text{Fe}^+$ implantation

Low Tc TES fabrication:

- CDMS W-TES through ion implantation, cryogenic testing
- CRESST W-alpha phase TES
- Or we can utilize superconducting bilayers as TES (proximity effect)

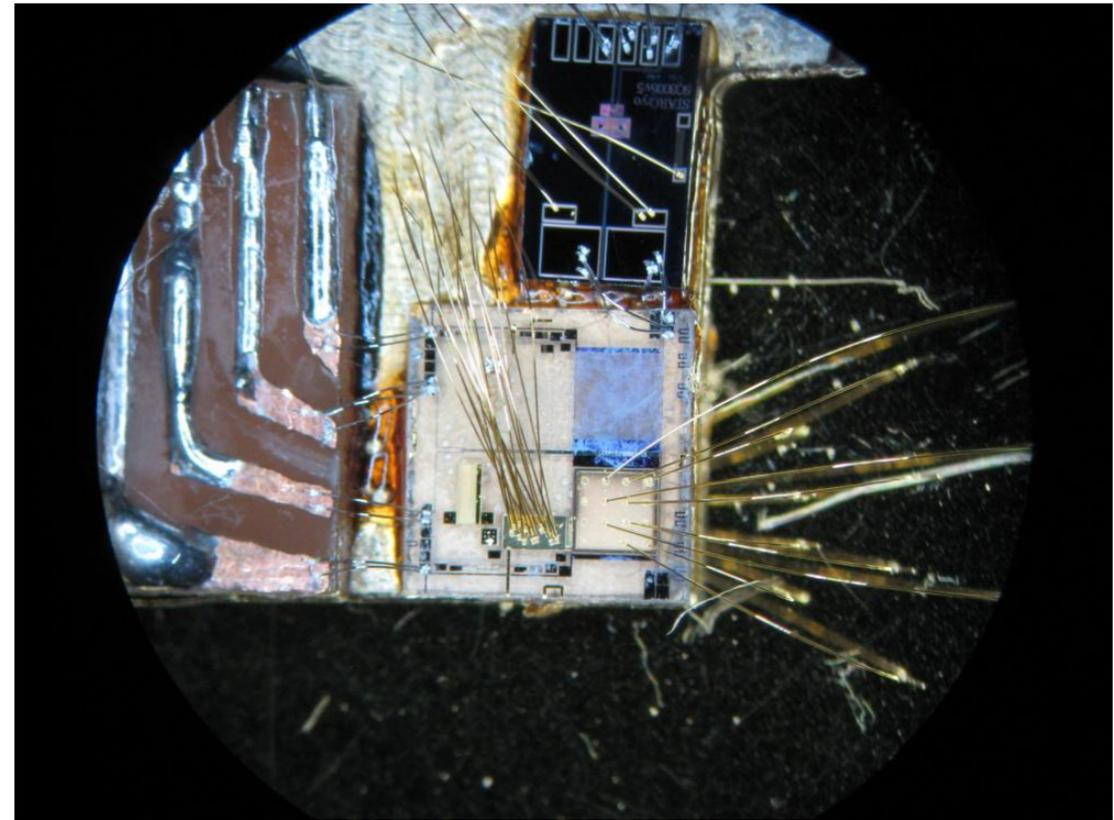
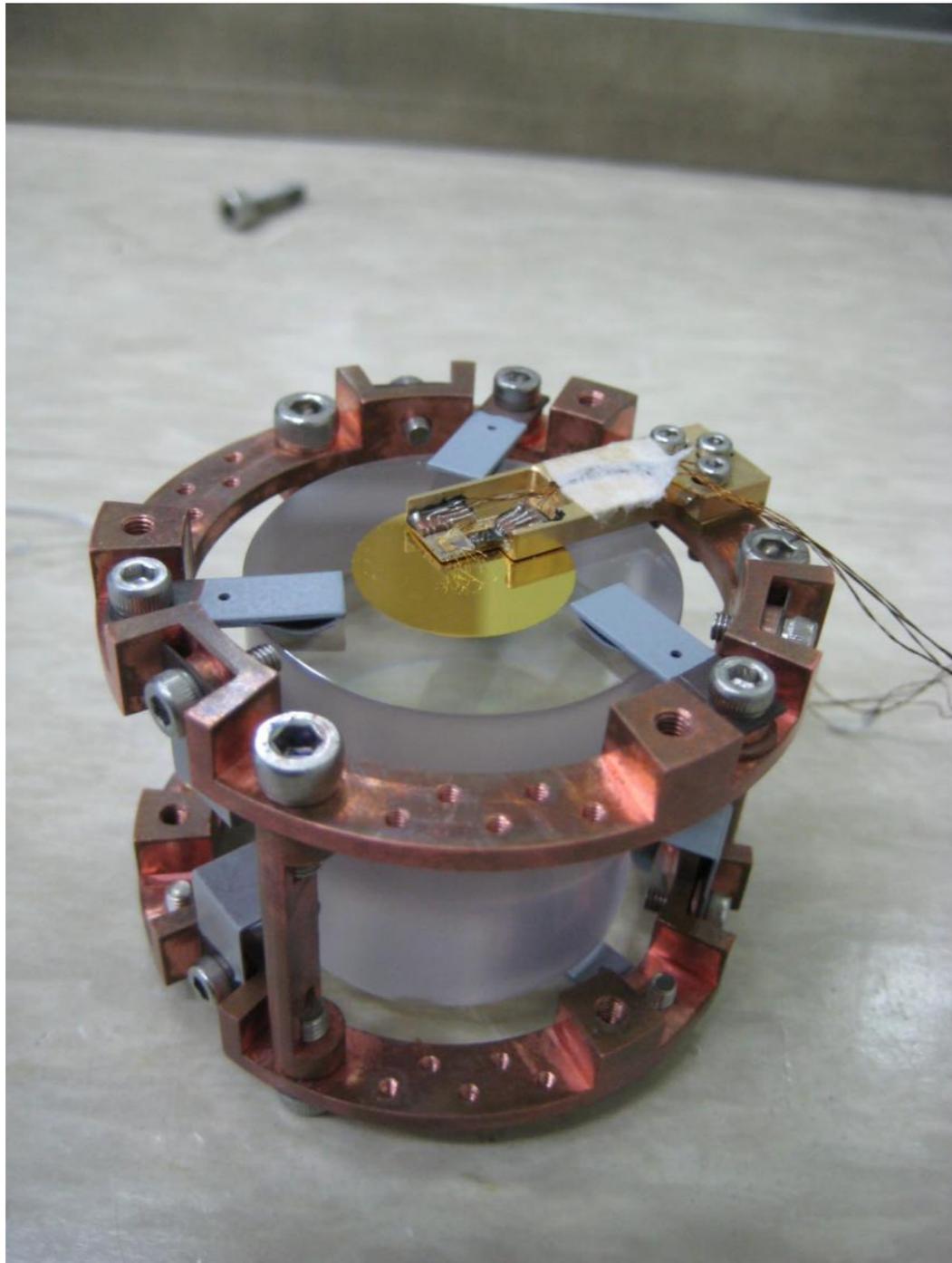
Large quantity (hundreds) production:

- Minimize # TES and SQUIDs on detector
- Simplify and tune thermal coupling TES/crystal-absorber
- Could the TES sit on the thermal bath?
- SQUIDs can be readout in arrays as large as up to 10,000

W-alpha phase for low Tc



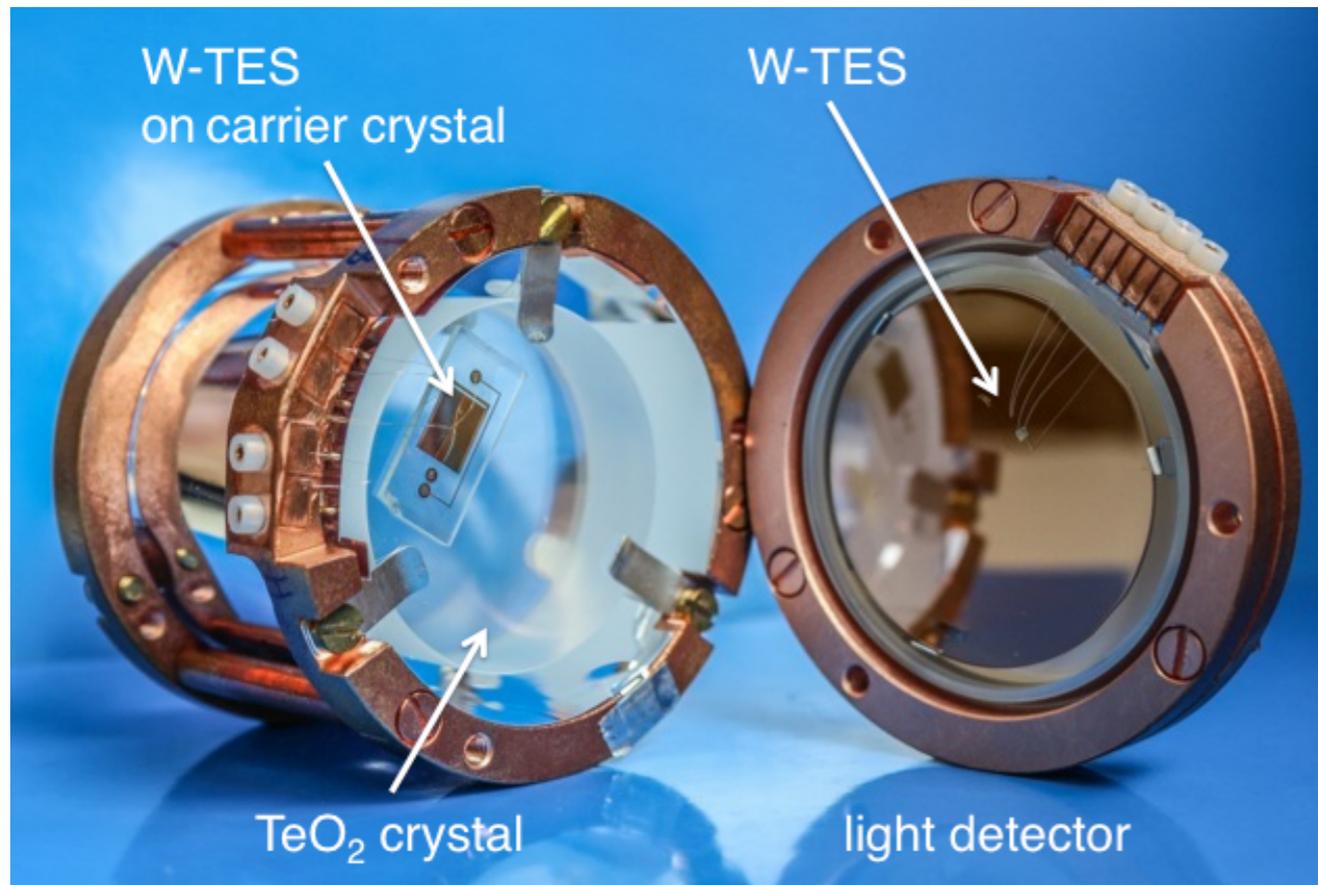
AMORE prototype from 2012: Au patch coupled via gold wires to a floating MMC - *scalability*



- 2cm diameter, 200nm thickness gold film is evaporated on the center of the crystal top surface.
- Meander chip is placed on the brass support and it is connected to the gold film by ~10 gold wires of 25um diameter.

For CUPID we need O(1000) channels

CRESST-II technology as Light Detector for TeO₂



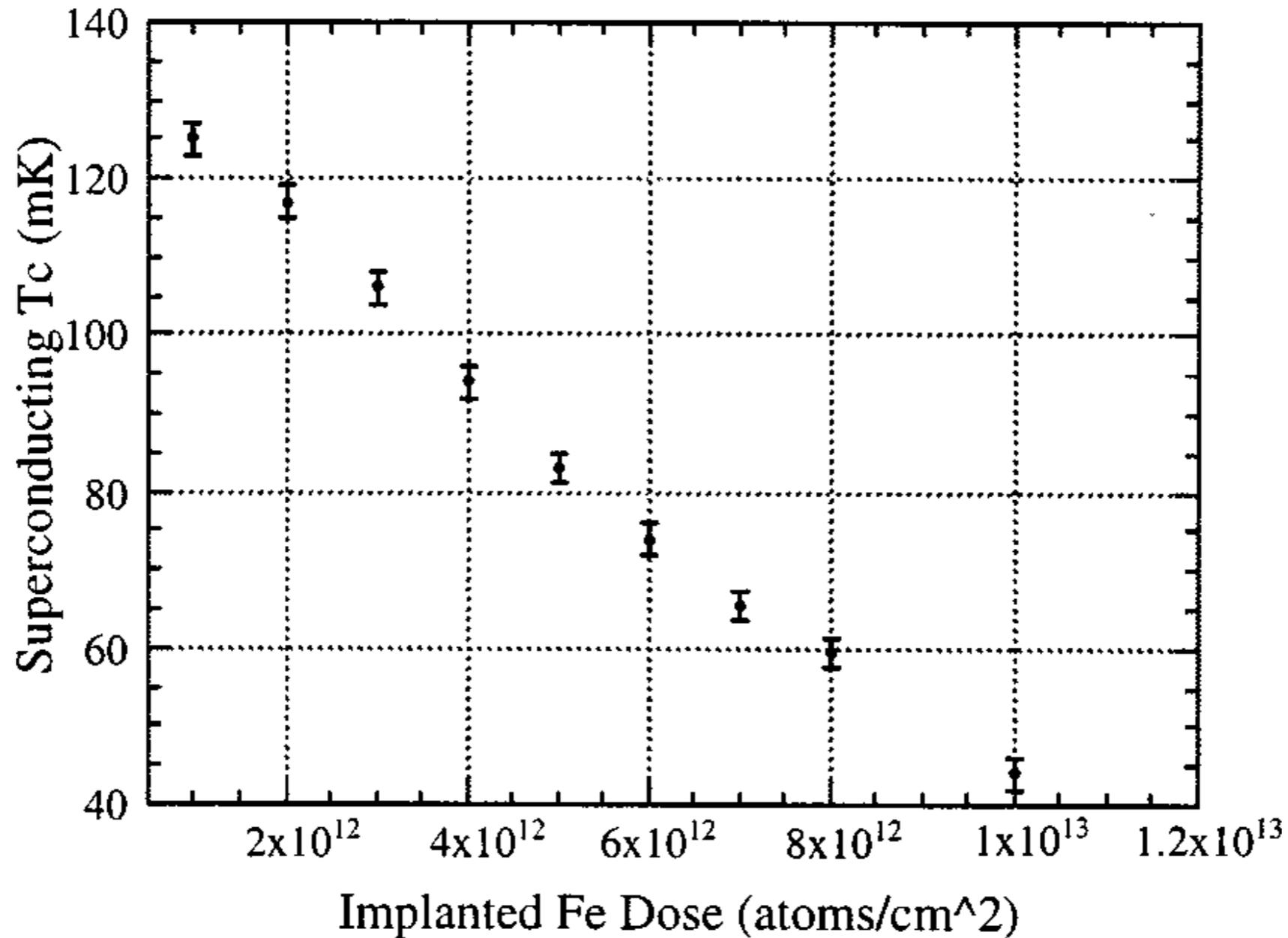
K. Schaffner et al. Astroparticle Physics (2015), pp. 30-36

- 280g TeO₂ crystal, 40mm diameter and height
- A small (20x10x2)mm³ CdWO₄ carrier crystal, equipped with the W-TES was attached on one of the flat surfaces of the TeO₂ by vacuum grease
- Light Detector CRESST-II type, sapphire disc (Silicon on Sapphire) readout by W-TES
- Light Detector reached RMS baseline $\sigma=24\text{eV}$, although CRESST-II has obtained $\sigma=5\text{eV}$

Absorber signal and sensor bandwidth mismatch minimization may improve energy resolution in this type of detectors as predicted by model in Pyle et al. arXiv:1503.01200

Tc tuning in W for TES applications in CDMS

J. Appl. Phys., Vol. 86, No. 12, 15 December 1999



Measured superconducting transition temperature for 350 Å-thick W films implanted with Fe-56 ions at 50 keV kinetic energy

Tc calculation for a superconducting bilayer using Usadel model

NIM A Vol. 444 (2000) 23-27

For thin films:

$$T_C = T_{C0} \left[\frac{d_s}{d_0} \frac{1}{1.13(1 + 1/\alpha)} \frac{1}{t} \right]^\alpha$$

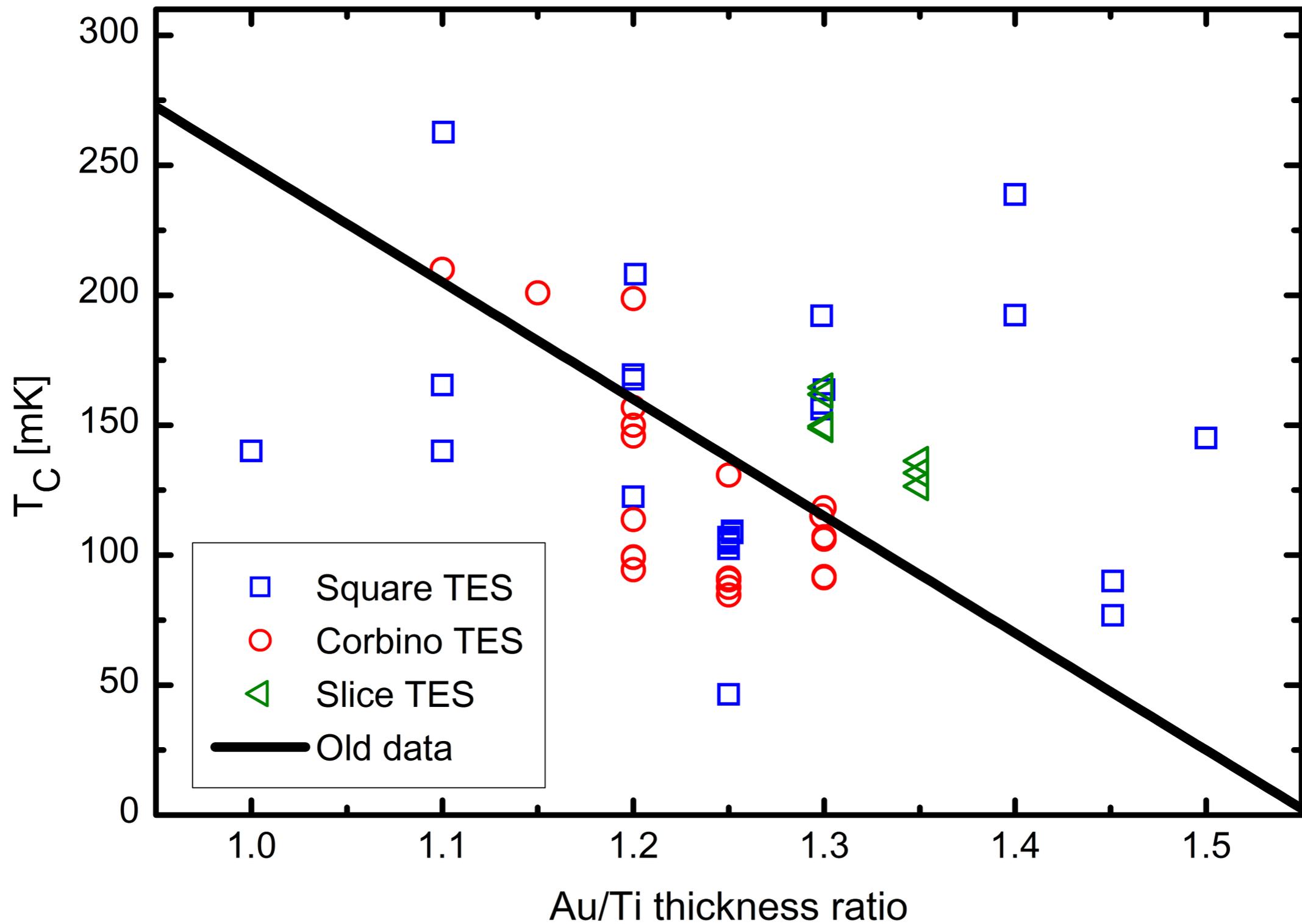
$$1/d_0 = (\pi/2)k_B T_{C0} \lambda_f^2 n_s$$

$$\alpha = d_n n_n / d_s n_s.$$

For a MoCu bilayer, $d_0 = 1.18 \mu\text{m}$, $n_n/n_s = 0.431$, $\lambda_f = 0.464\text{nm}$.
From data they obtained $t = 0.21$

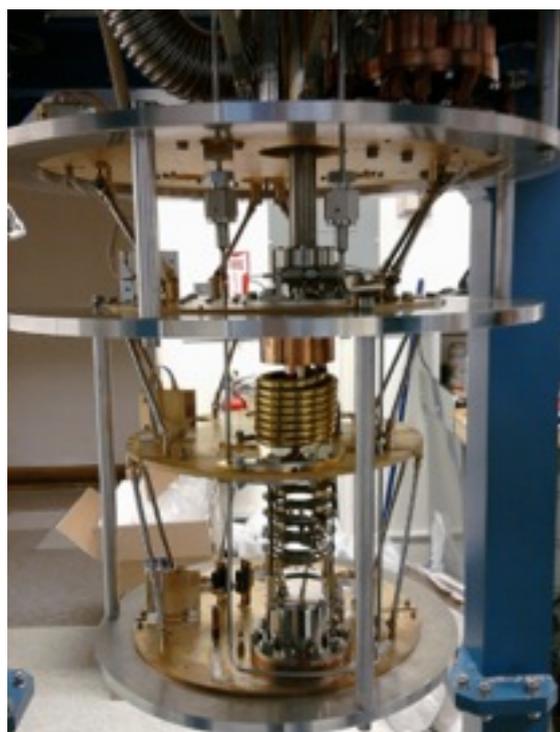
Although experimentally is a very challenging task

Kinnunen PhD Thesis

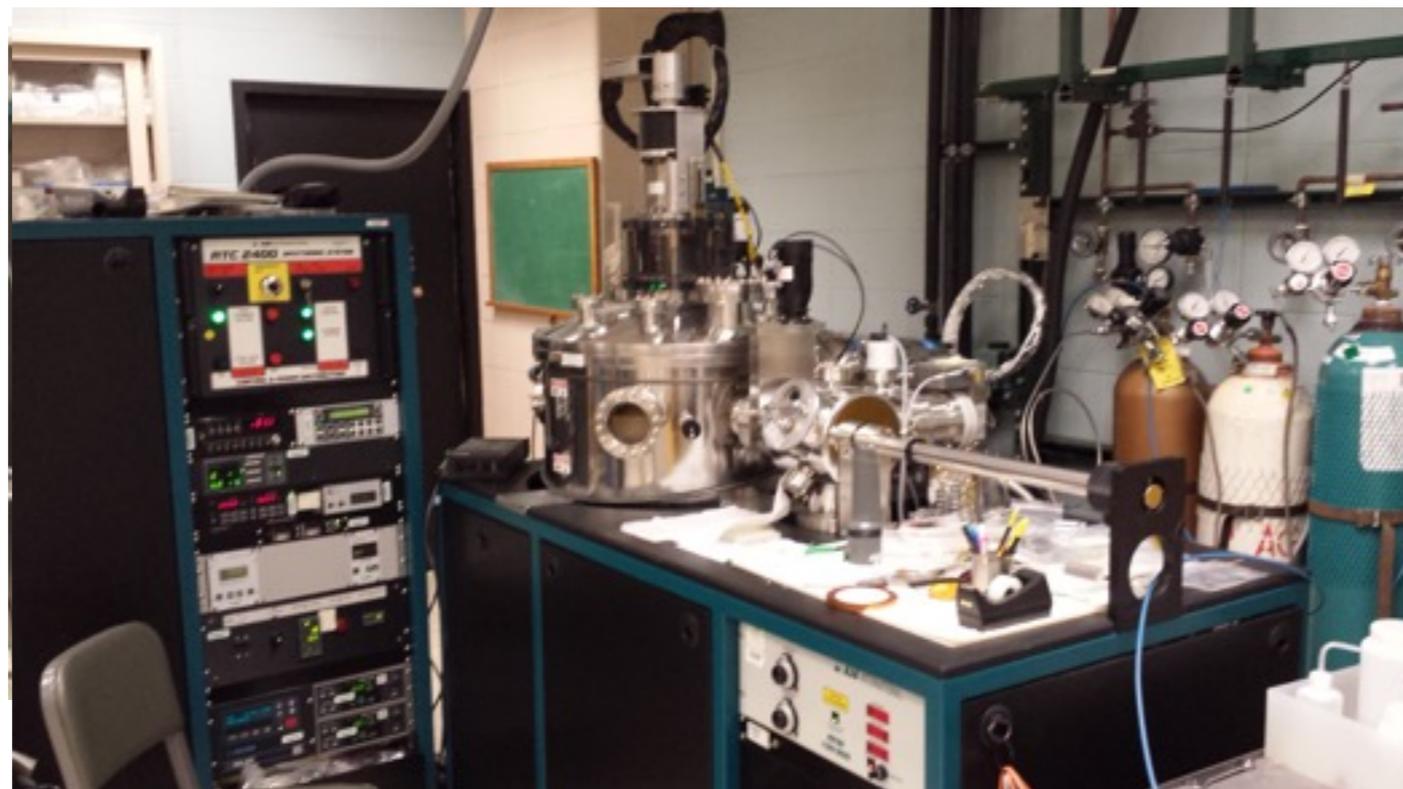


Sample production at ANL and testing at Berkeley

Collaborating at ANL with G. Wang, V. Novosad, V. Yefremenko and C. Chang



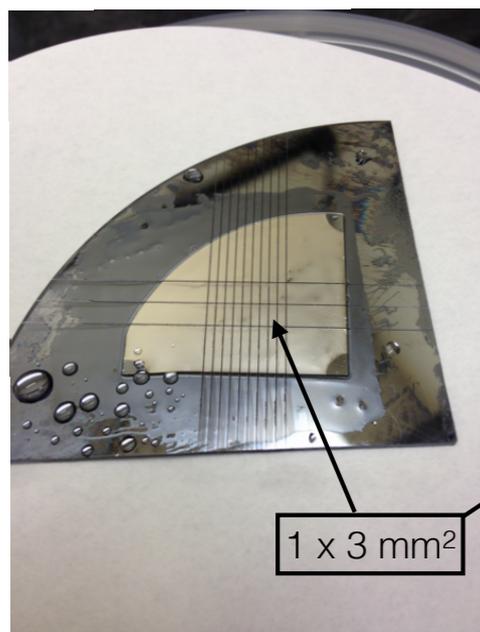
Cryogen free DR and Co-60 thermometry down to $\sim 7\text{mK}$



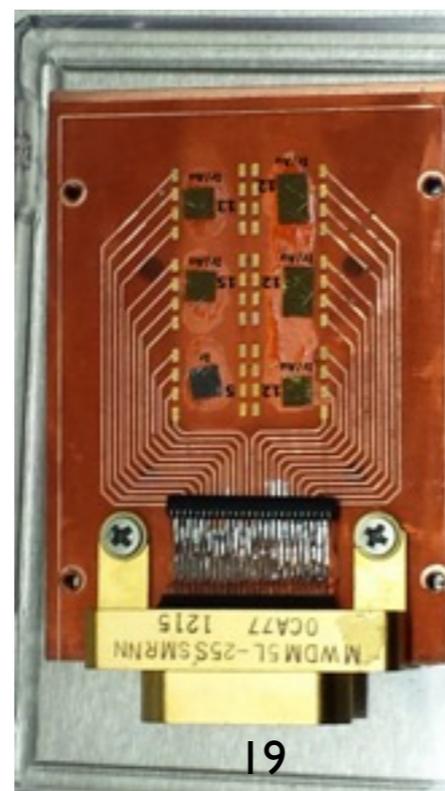
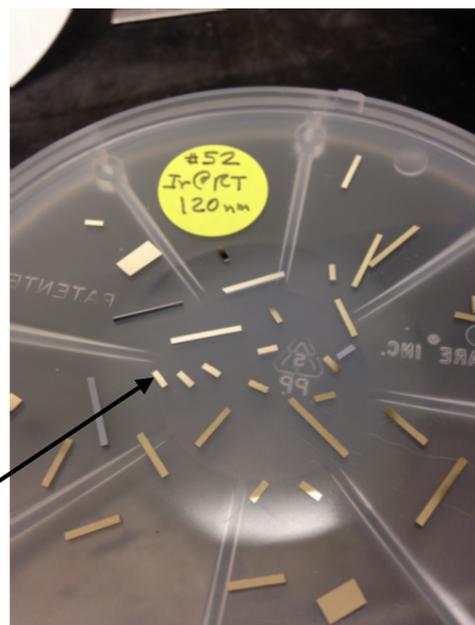
At ANL with sputtering chamber on the back



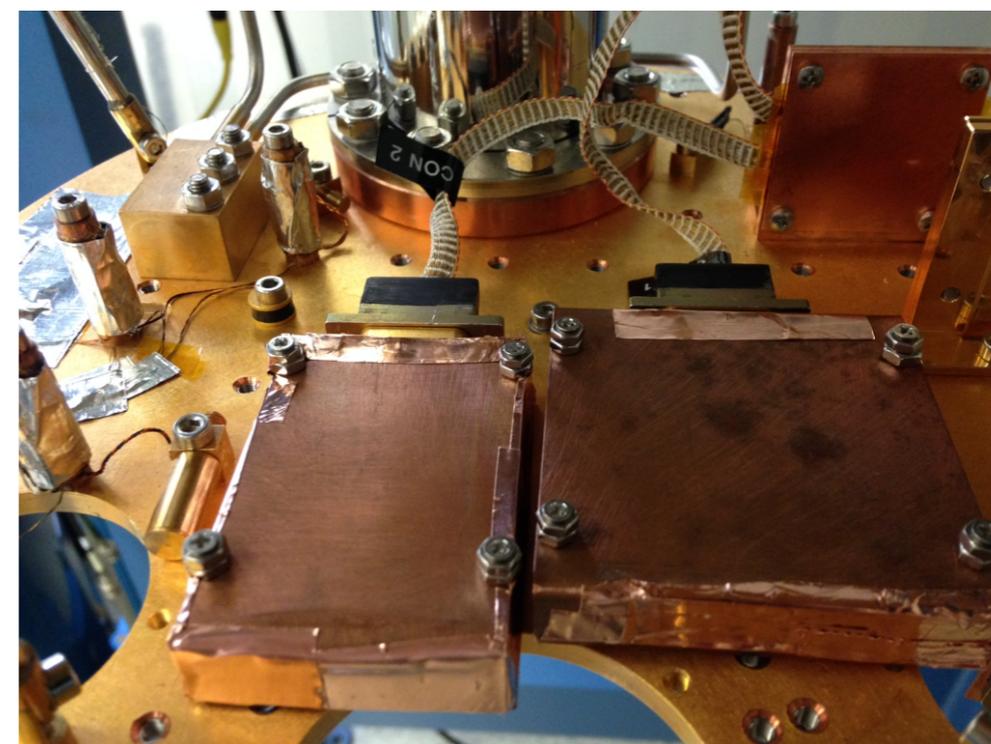
Iridium bilayers



$1 \times 3 \text{ mm}^2$

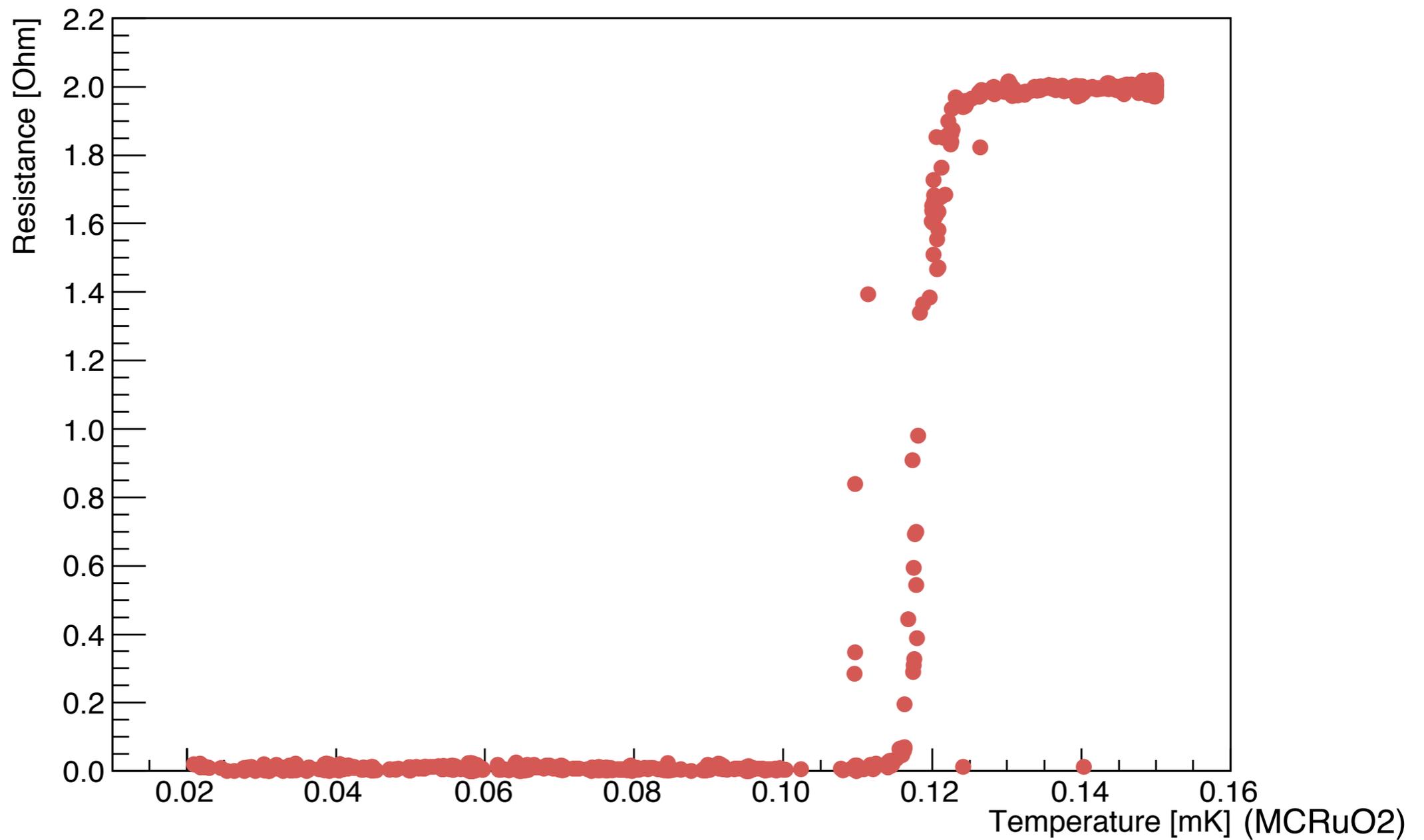


19



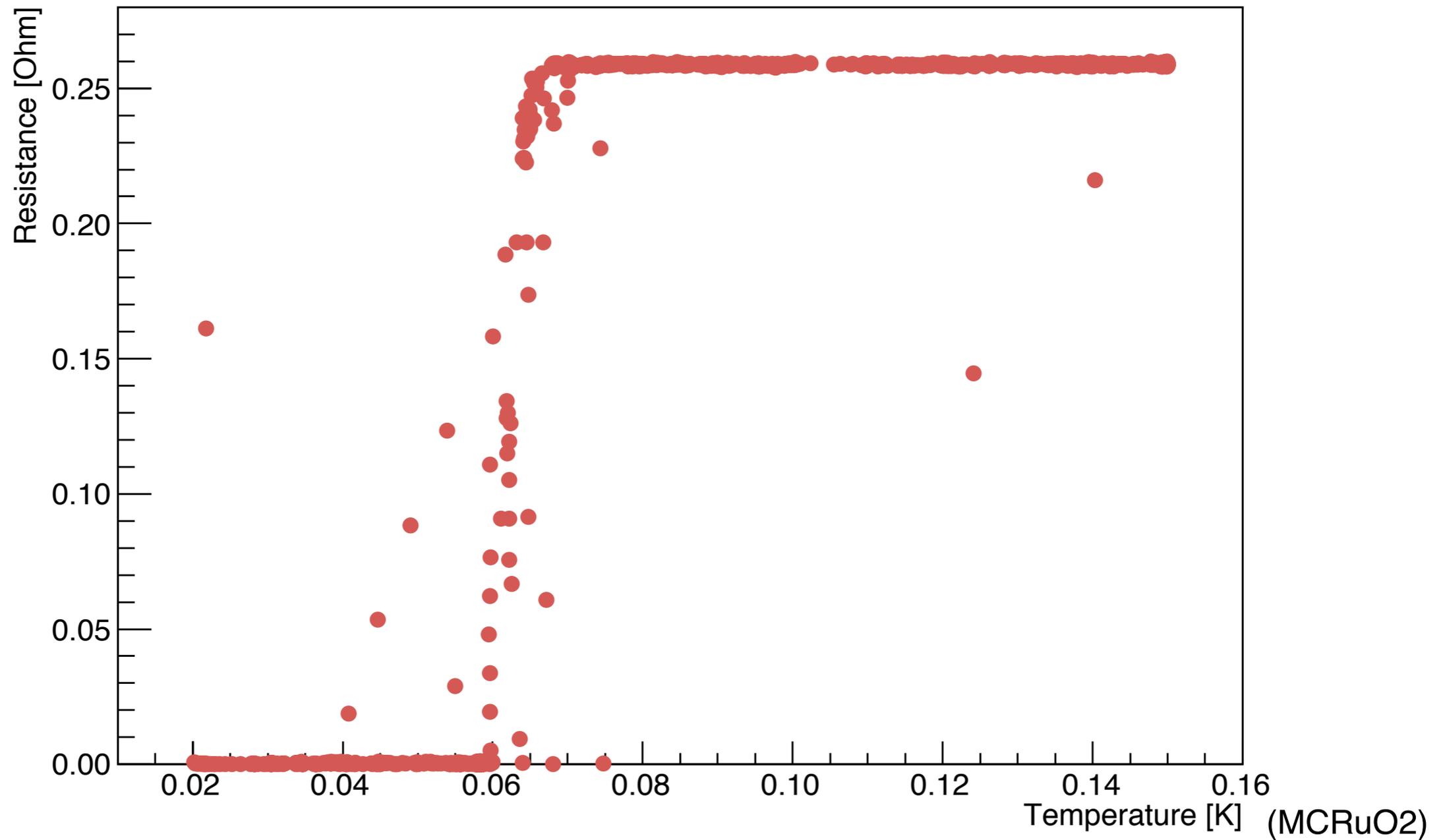
Ir at RT / Au
Ir = 80 nm
Au = 120 nm

$T_c = 120$ mK
Channel 07



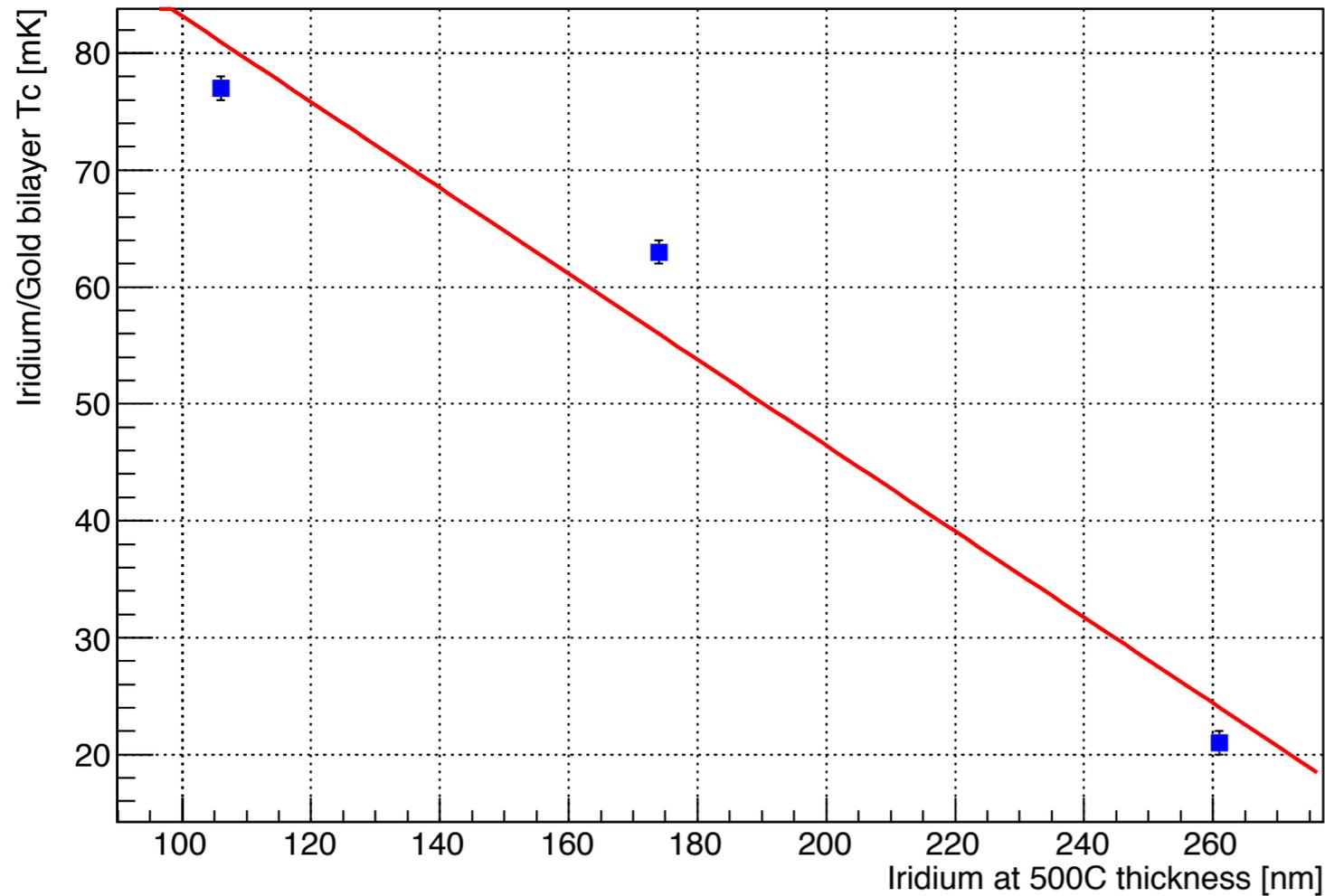
Ir at 600 °C / Au
Ir = 80 nm
Au = 120 nm

$T_c = 60$ mK
Channel 06

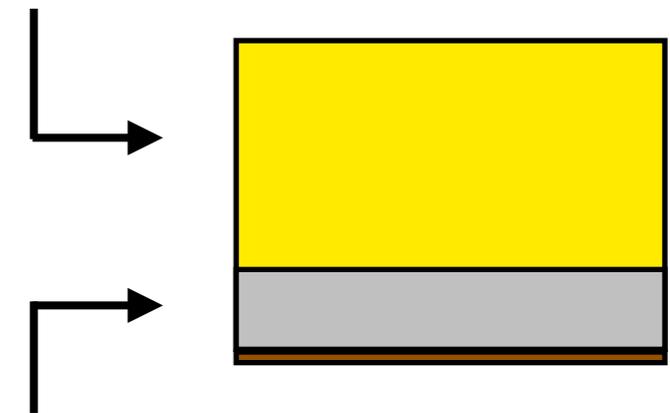


Ir@500C/Au bilayer T_c as a function of Au thickness

Ir = 101nm at 500C / Au = x-axis thickness



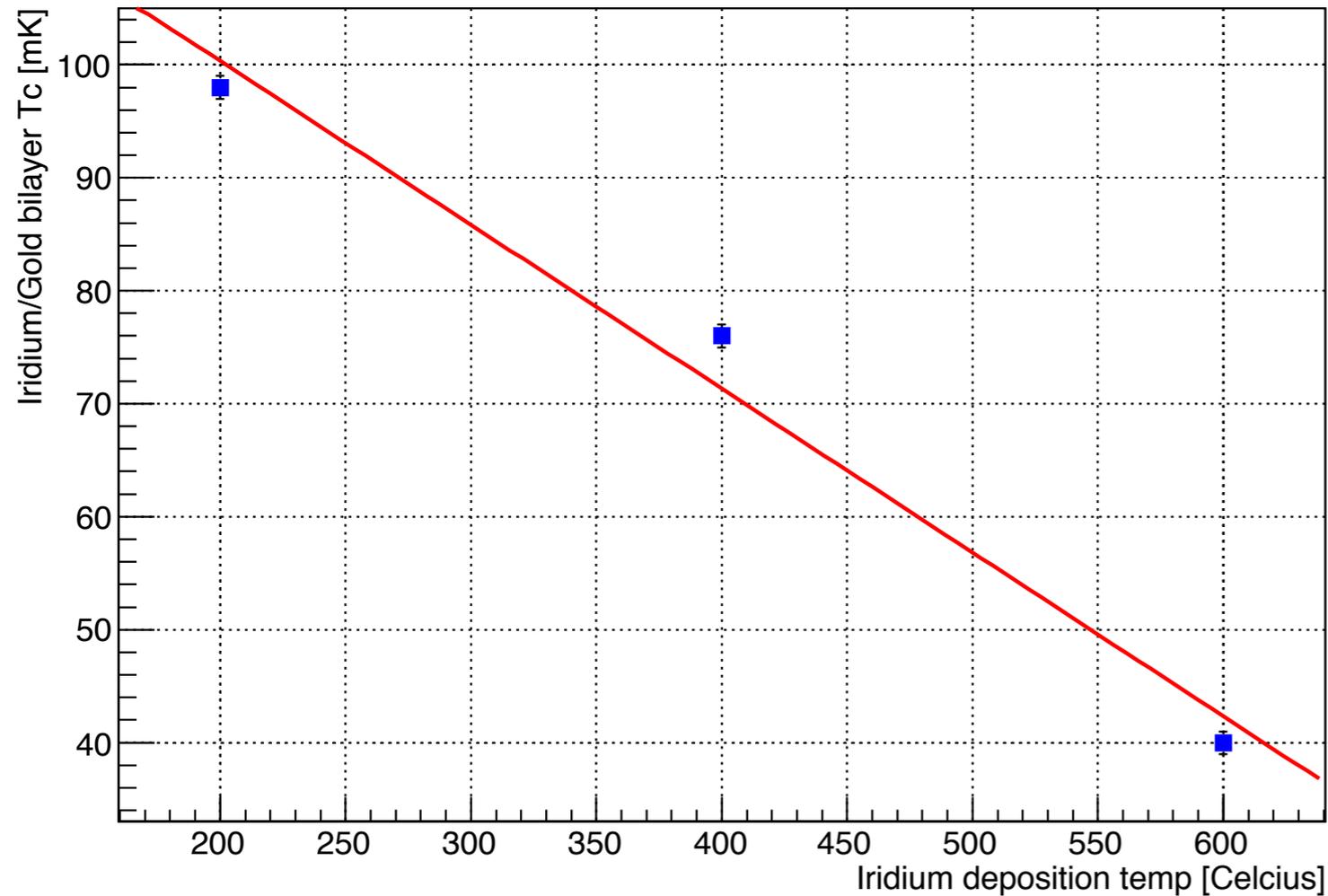
Au layer at room temp



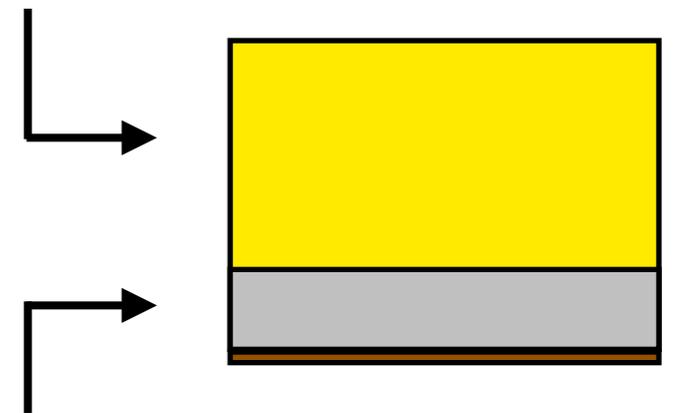
100nm Ir deposited at 500C

Ir/Au bilayer T_c as a function of Ir temp

Ir = 80nm at x-axis temp / Au = 160nm



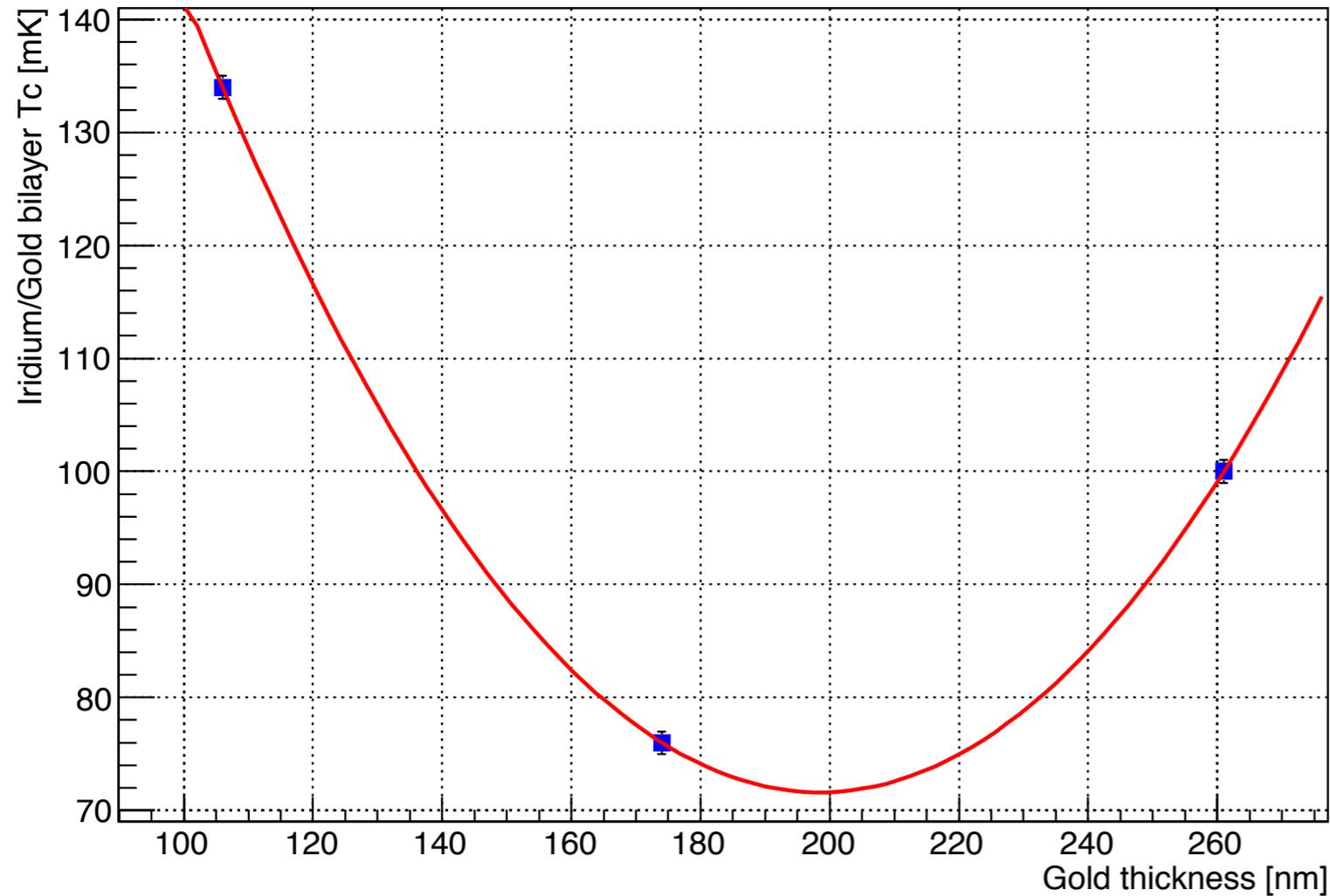
260nm Au at room temp



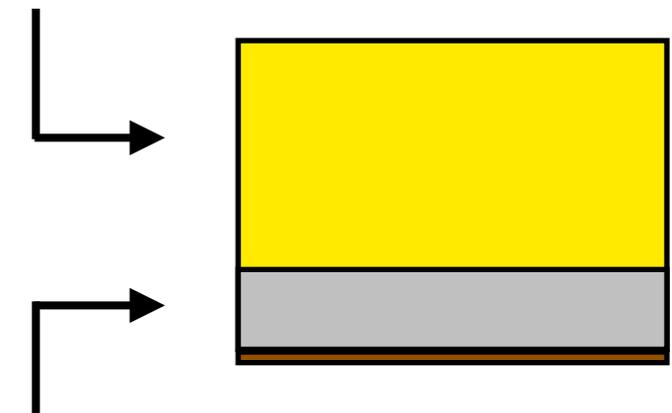
100nm Ir annealed at different temperatures

Ir/Au at Room Temperature was discarded

Ir = 101nm at room temp / Au = x-axis nm

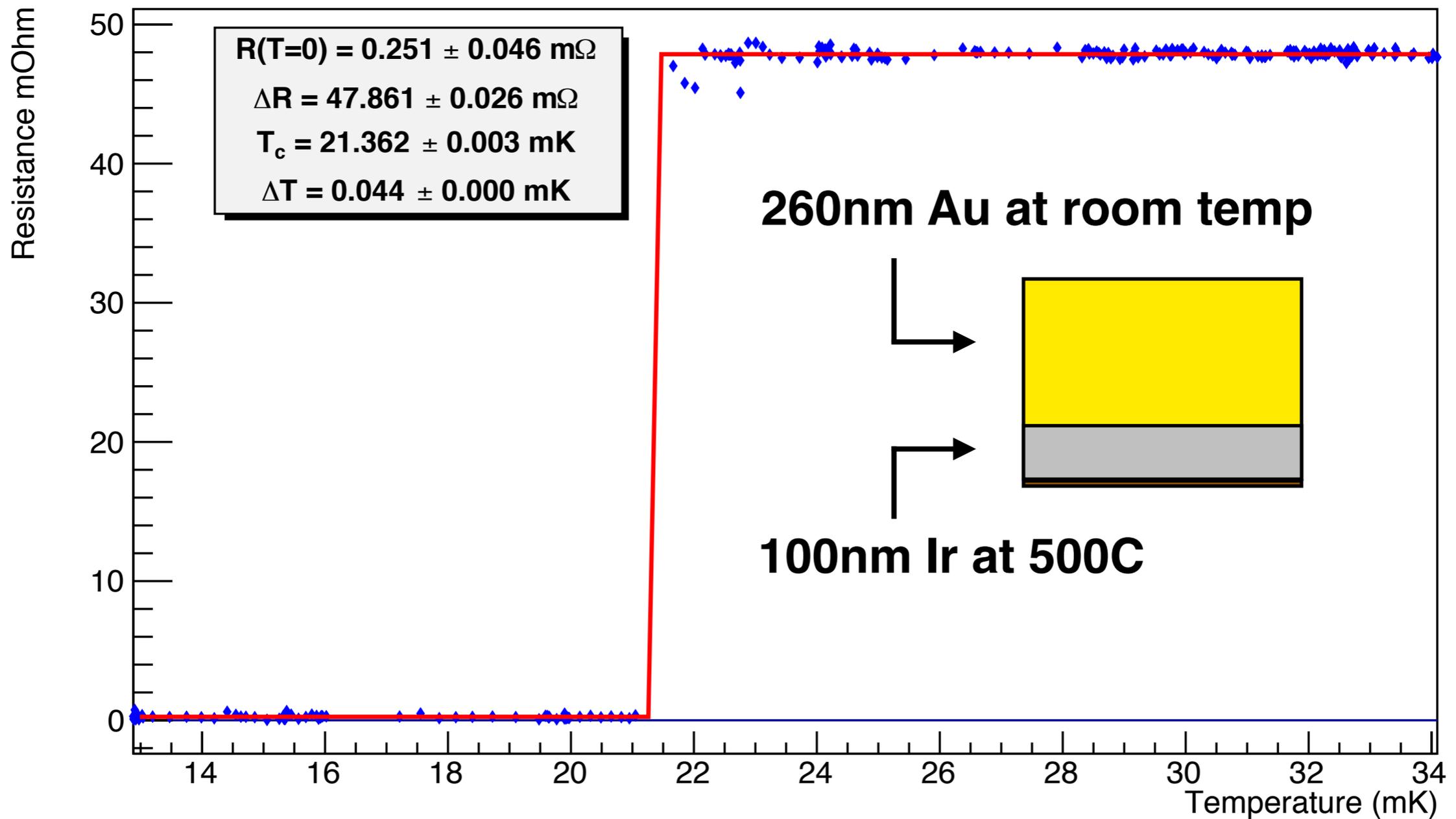


100-260nm Gold



100nm Ir at room temp

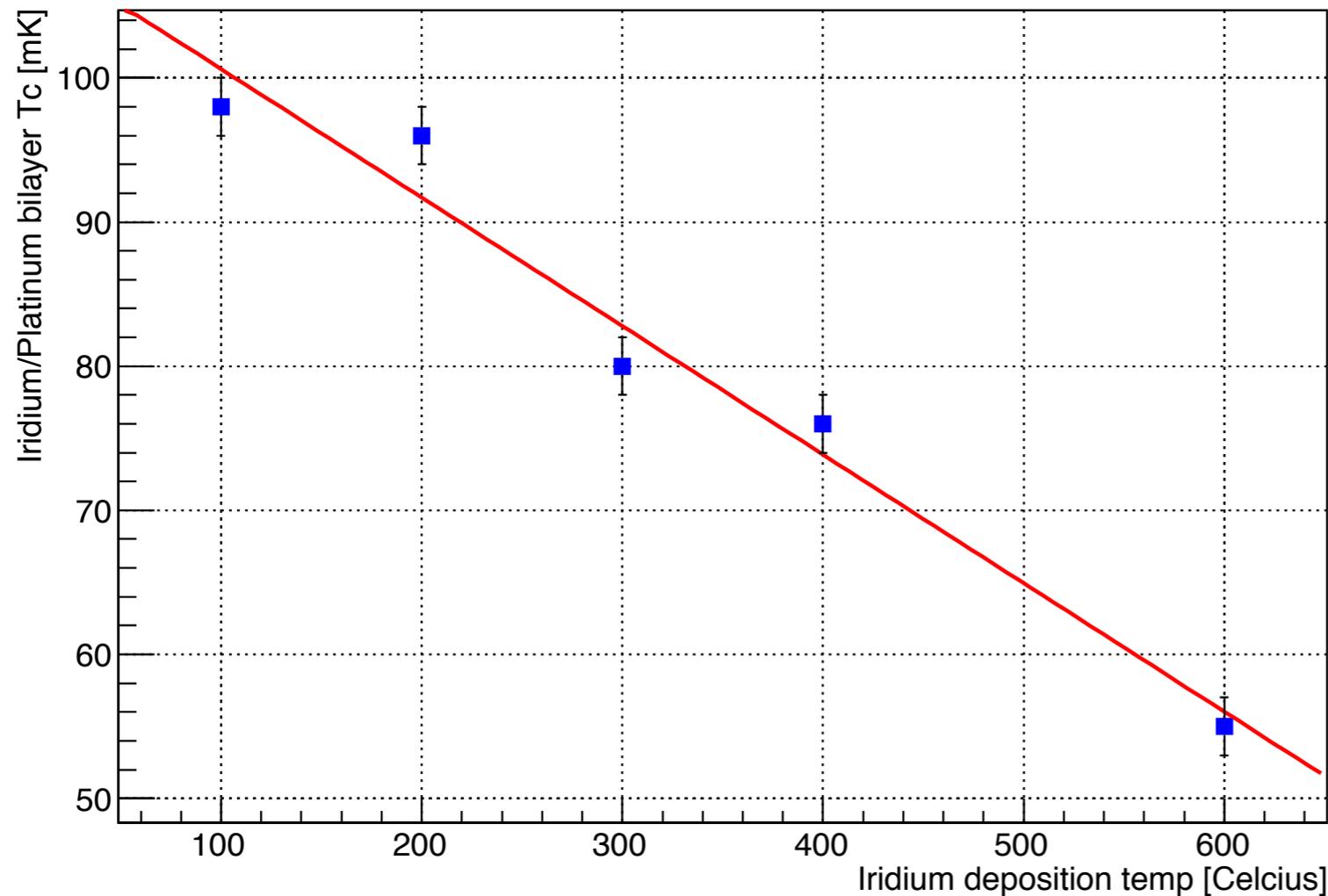
Our lowest Ir/Au bilayer $T_c=21.4$ mK



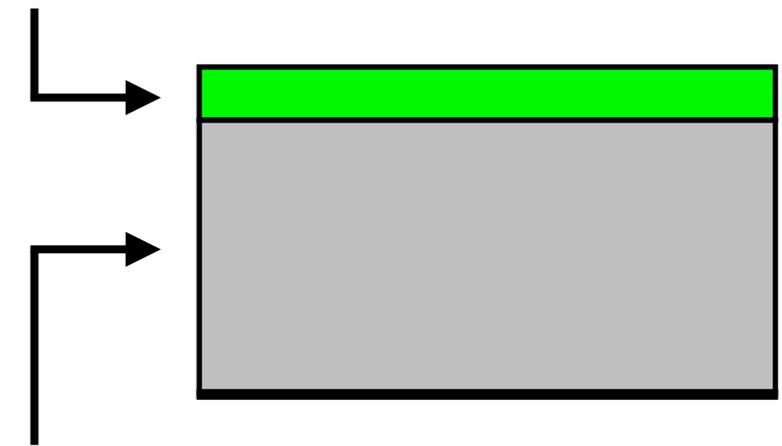
A promising new bilayer for Low Tc TES: Ir/Pt

Tc dependence on Iridium deposition temperature for an Ir/Pt

Ir = 80nm at x-axis temp / Pt = 20



20nm Pt at room temp



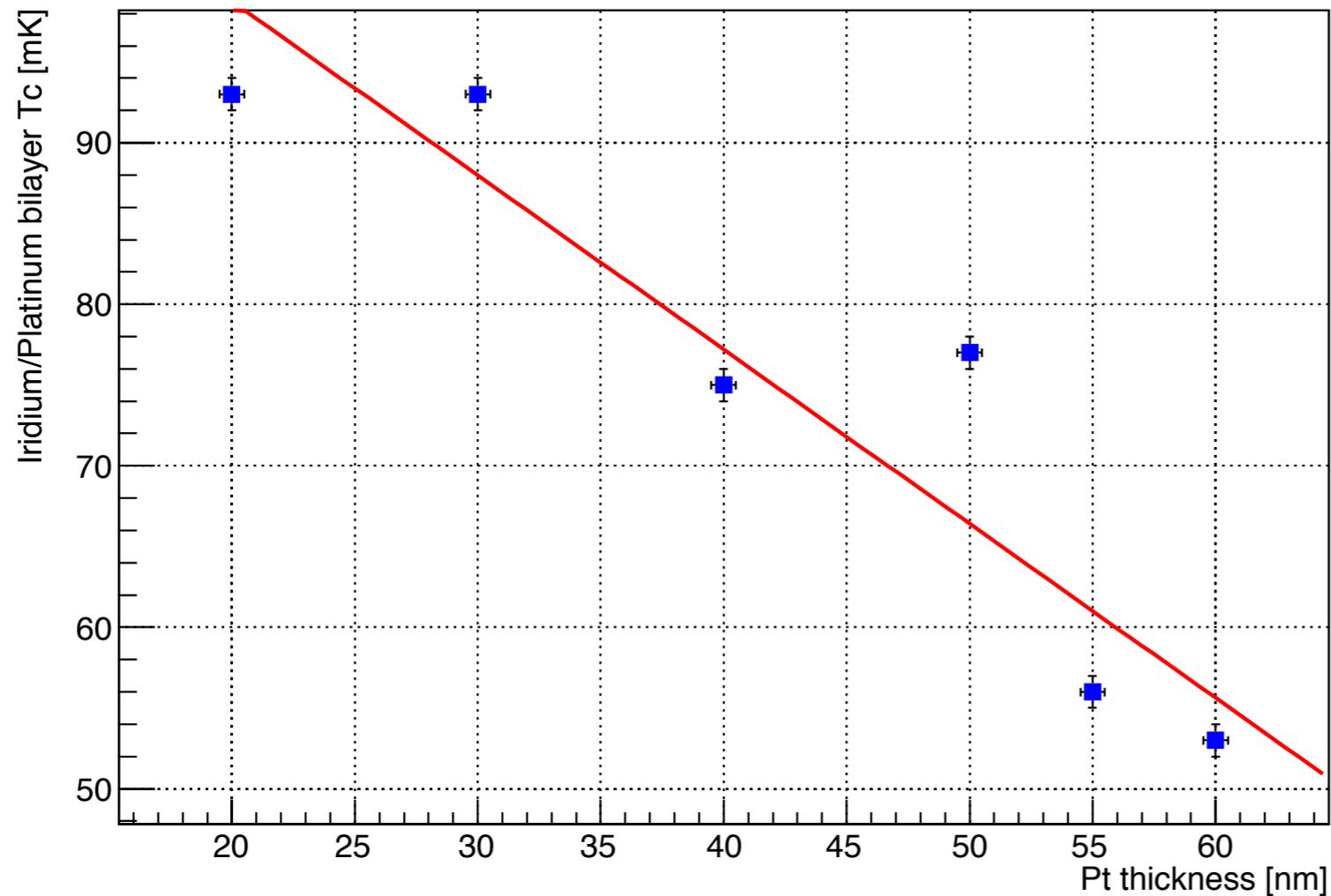
80nm Ir at varying temp

Normal resistance $\sim 10x$ of IrAu for same Tc

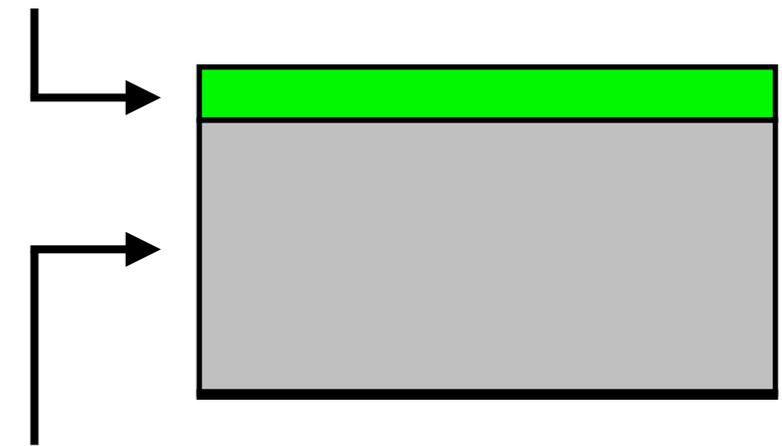
Ir/Pt at room temperature

Tc dependence on Platinum thickness

Ir = 100nm at Room Temp / Pt = x-axis



20nm Pt at room temp



80nm Ir at varying temp

Room temperature deposition of TES would make application to large crystals possible

Summary

- TES technology for Double Beta Decay may be applicable in order to improve energy resolution and reduce backgrounds
- TES as light detectors for macro-bolometers in large quantities $O(1000)$ remains to be demonstrated for the future of double beta decay with bolometers
- Low T_c material for TES with fast fabrication time may be achievable utilizing Iridium/Gold and Iridium/Platinum
- Ir/Au only works if we heat Ir and let it anneal to room temperature before depositing Au at room temperature
- We have found a new promising bilayer, Ir/Pt that works both with Ir annealing and at room temperature. It also shows large $R(\text{normal})$ compared to Ir/Au ($\sim 10x$).
- Large T_c suppression (reproducibility) remains an issue without dedicated film deposition chambers
- SQUID multiplexing may be required, although the channel number has been already accomplished by other experiments in CMB community